

INTRODUCTION

Olivine crystals in andesites, basalts, picritic melts, komatiites and mantle peridotite have Ni contents that range from about 1000 to 5000 ppm (Sobolev et al. 2007; Straub et al. 2008; Herzberg et al. 2013). This variability is a mineralogical record of magmatic origins. Nickel is no different from any other element in having abundance levels that depend on how it is partitioned amongst phases, processes such as mixing and melt-rock reaction, T-P conditions of partial crystallization and melting, in addition to the Ni contents of the sources that undergo melting. But what distinguishes nickel is that it is concentrated in olivine relative to the magmas from which it crystallizes (Hart and Davis 1978; Beattie et al. 1991; Matzen et al. 2013), and there is an abundance of data because it is easy to analyze. The problem is how to reliably extract information about origin from the Ni contents of olivine. We show that olivine chemistry provides a remarkable mineralogical narrative about magmatic processes in Earth's crust, mantle, and possibly core over the entire age of the Earth.

The narrative that olivine provides relies heavily on understanding how Ni is partitioned between olivine and liquid (i.e., $D_{\text{Ni}}^{\text{Ol/L}}$), which depends on temperature, pressure, and composition of the melt (Hart and Davis 1978; Beattie et al. 1991; Wang and Gaetani 2008; Filiberto et al. 2009; Li and Ripley 2010; Putirka et al. 2011; Niu et al. 2011; Matzen et al. 2013). $D_{\text{Ni}}^{\text{Ol/L}}$ decreases with increasing MgO content of the melt. But as increasing temperatures and pressures of melting yield melts with higher MgO contents (O'Hara 1968; Herzberg et al. 2007), it can be difficult to resolve the separate temperature (T)-pressure (P)- and composition (X) effects. This problem has given rise to a plethora of parameterization models of experimental data that calibrate $D_{\text{Ni}}^{\text{Ol/L}}$ as a function of T-P-X. It is an important problem to resolve because elevated Ni contents of olivine phenocrysts have been used to infer either pyroxenite melting in the source (Sobolev et al. 2005; 2007; Herzberg 2011), or elevated temperatures and pressures of melting of a peridotite source (Li and Ripley 2010; Niu et al. 2011; Putirka et al. 2011; Matzen et al. 2013).

We assume that nickel is partitioned between olivine and liquid as described by the empirical Beattie-Jones model, first formulated by Jones (1984) and later calibrated by Beattie et al. (1991). This model requires no independently adjustable temperature and pressure terms to understand how Ni is partitioned between olivine and melt. We discuss the advantages and limitations of this model, and compare it with the model of Matzen et al. (2013), which has the advantage that it is a thermodynamic parameterization of experimental data. The Beattie-Jones model is used in conjunction with an Fe-Mg partitioning parameterization of experimental data for olivine and melt (Toplis 2005; Herzberg and O'Hara 2002) to fully describe a forward model of the compositions of peridotite partial melts and the olivines from which they crystallize. It is shown that this olivine model provides an excellent description of measured olivines with relatively low Ni contents (i.e., $< \sim 3500$ ppm), and it can deepen our understanding of the origin of mantle peridotite, Archean komatiites, and Phanerozoic magmatism from oceanic spreading centers and mantle plumes.

COMPUTATIONAL METHOD

The method for calculating model olivine Ni compositions was given elsewhere (Herzberg and O'Hara 2002; Herzberg 2011) and is here developed and extended to include cobalt. Primary magma compositions are calculated first, followed by the olivines that they crystallize. We also examine the effects of fractional crystallization on modifying primary magmas, and provide examples of olivines that are expected to crystallize along a liquid line of descent; such olivine compositions have been appropriately termed "crystal line of descent", or CLD (Putirka et al. 2011).

The compositions of Ni in liquids extracted from dunite [L + Ol] and harzburgite [L + Ol + Opx] residues are obtained from mass balance solutions to the equation for accumulated fractional melting:

$$C_L = C_o [1 - (1-F)^{1/D}] / F \quad (1)$$

where C_L = weight % NiO in the liquid (primary magma), C_o = initial NiO in the peridotite source composition, F = melt fraction, and D is the bulk distribution coefficient. This is suitable for most geological occurrences because melts drain by buoyant porous flow from their sources at low melt fractions (Ahern and Turcotte 1979; McKenzie 1984). However, we also consider batch melting equation:

$$C_L = C_o / (F(1-D) + D) \quad (2)$$

which is appropriate for understanding high pressure experiments on KLB-1 (Herzberg and Zhang 1996), and we argue that it may have been important in the crystallization of Earth's magma ocean.

The Beattie-Jones model has been assumed for computing the compositions of olivine in equilibrium with a liquid of a known composition. The liquid composition in weight % oxides is first recalculated to mole % oxides on a mono-cation basis (i.e., MgO, NiO, $AlO_{1.5}$, $NaO_{0.5}$, etc), and the compositions of olivine are computed using the molar percent partition coefficients:

$$D_i^{Ol/L} = X_i^{Ol} / X_i^L \quad (3)$$

where X_i refers to the mole fraction of oxide component i in liquid (L) and olivine (Ol). And in equations (1) and (2), the bulk distribution coefficient D is equal to $D_i^{Ol/L}$ and $D_i^{Opx/L}$ weighted according to the mass fractions of olivine and orthopyroxene that participate in melting, where $D_{NiO}^{Opx/L}$ and $D_{CoO}^{Opx/L}$ are from Beattie et al. (1991).

In the Beattie-Jones model, D_i depends on the partitioning of MgO between mineral and liquid. Based on an existing database at the time, Beattie et al. (1991) parameterized the experiments with the simple linear equations:

$$D_{NiO}^{Ol/L} = 3.346 D_{MgO}^{Ol/L} - 3.665 \quad (4)$$

130

131 and:

132

$$133 \quad D_{\text{CoO}}^{\text{Ol/L}} = 0.786 D_{\text{MgO}}^{\text{Ol/L}} - 0.385 \quad (5)$$

134

135 In equations (4) and (5):

136

$$137 \quad D_{\text{MgO}}^{\text{Ol}} = \text{MgO}^{\text{Ol}}/\text{MgO}^{\text{L}} \quad (6)$$

138

139 In these equations, the partitioning of Ni depends critically on the MgO content of the
140 magma, which we explore from solidus to liquidus melting conditions for fertile peridotite.
141 Experimental data of Walter (1998) on peridotite source KR4003 were used for much of the
142 petrological modeling (Herzberg and O'Hara 2002). KR4003 contains 38.12% MgO (Herzberg
143 and O'Hara 2002; Herzberg and Asimow 2015), slightly more depleted than the model pyrolitic
144 mantle of McDonough and Sun (1995). Although KR4003 was reported to contain 0.24% NiO
145 (Walter 1998), we assume that it contains the canonical value of 0.25% NiO (i.e., 1964 ppm;
146 McDonough and Sun 1995). The major element geochemistry of model primary magmas for
147 KR4004 has been given previously (Herzberg and O'Hara 2002; Herzberg 2004; Herzberg and
148 Asimow 2015).

149

150 It is quite remarkable that the Beattie et al. (1991) parameterization performs extremely
151 well in describing more recently acquired experimental data discussed below. However, we
152 have chosen not to update the parameterized constants in equations (4) and (5) because the Ni
153 and Co contents of computed olivines and liquids are not greatly improved, and they are well
154 within the uncertainties discussed below.

155

156 The Ni contents for near-solidus melts were calculated from the MgO contents of near-
157 solidus melts (Herzberg and O'Hara 2002) together with a solidus olivine composition having
158 0.365% NiO, which is characteristic of high precision analysis of olivines having an Mg number
159 of 89.5 (Herzberg et al. 2013; see below). In a similar way, we constrain the Co content of near-
160 solidus melts using olivine with 0.0172% CoO as we report below for high precision electron
161 microprobe analysis of San Carlos olivine. It will be shown below that Ni contents of solidus
162 melts are similar to or slightly higher than those for L+Ol and L+Ol+Opx, placing confidence in
163 all interpolation possibilities for spinel and garnet peridotite.

164

165 The Mg number (i.e., $100\text{MgO}/(\text{MgO} + \text{FeO})$; cationic) of olivine is computed using the
166 partitioning of Fe and Mg between olivine and liquid:

167

$$168 \quad K_{\text{D}}^{\text{Ol/L}}_{\text{FeO/MgO}} = D_{\text{FeO}}^{\text{Ol/L}} / D_{\text{MgO}}^{\text{Ol/L}} \quad (7)$$

169

170 For crystallization of olivine at near-surface conditions, we use the thermodynamic
171 parameterization of Toplis (2005), wherein K_{D} is typically ~ 0.30 for mid ocean ridge-like
172 basalts, as originally constrained by Roeder and Emslie (1970). However, K_{D} is elevated at high
173 pressures in the melting region, and in this case we use the parameterization of Herzberg and
174 O'Hara (2002) for batch and accumulated fractional melting; application of this model to the

175 high pressure experimental results of Walter (1998) yields K_D that is very similar to that obtained
176 from the Toplis (2005) model.

177

178

179

COMPUTATIONAL UNCERTAINTIES

180

181 The advantage of the Beattie-Jones model is that it requires no independently adjustable
182 temperature and pressure terms to understand how Ni is partitioned between olivine and melt;
183 this arises because elevated temperatures and pressures of melting typically result in elevated
184 MgO contents of primary magmas of mantle peridotite, and the T-P variability is captured to a
185 large extent by the $D_{\text{MgO}}^{\text{Ol/L}}$ term. A disadvantage is that the Beattie-Jones model is not grounded
186 in a thermodynamic analysis, and there may be situations in nature wherein variations in $D_{\text{NiO}}^{\text{Ol/L}}$
187 are not adequately captured by $D_{\text{MgO}}^{\text{Ol/L}}$ in the parameterized experimental database. In contrast,
188 the model of Matzen et al. (2013) describes $D_{\text{Ni}}^{\text{Ol/L}}$ with independently adjustable T-P-X terms in
189 a Gibbs energy equation and, in principle, it should be a more robust method. However, a
190 disadvantage of the Matzen et al. (2013) model is that it requires an understanding of the T-P
191 conditions of melting and crystallization. While improved constraints are now becoming
192 available (e.g., Ghiorso and Sack 1995; Herzberg and Asimow 2015; Hole and Millet, 2016), the
193 absolute T and P melting conditions are models that add a new layer of uncertainty. And in the
194 case of magma ocean formation considered below, the T-P conditions of melting are not well
195 known, and it is under these circumstances where the Beattie-Johns model is particularly useful.

196

197 In a previous paper (Herzberg et al. 2014), it was suggested that the Beattie-Jones model
198 was more accurate than the Matzen et al. (2013) model because it provided the minimum
199 absolute root mean square error (A RMSE). This analysis was based on comparison of the two
200 different models based on the absolute difference between experimental and model values of
201 $D_{\text{NiO}}^{\text{Ol/L}}$ (i.e., $D_{\text{Ni}} \text{ XP} - D_{\text{Ni}} \text{ Model}$) so that the absolute mean square error is:

202

$$203 \quad \text{A RMSE} = (\Sigma(D_{\text{Ni}} \text{ XP} - D_{\text{Ni}} \text{ Model})^2/N)^{0.5} \quad (8)$$

204

205 where $D_{\text{Ni}} \text{ XP}$ is the experimentally measure value of $D_{\text{NiO}}^{\text{Ol/L}}$ and $D_{\text{Ni}} \text{ Model}$ is $D_{\text{NiO}}^{\text{Ol/L}}$
206 predicted by the model. The problem with this approach is that the Matzen et al. (2013) model
207 $D_{\text{NiO}}^{\text{Ol/L}}$ values as expressed on a molar basis are higher for individual experiments than those of
208 Beattie-Jones, which gives rise to an artificially high A RMSE. A better way to compare
209 partitioning different models is to use the *relative* difference between experimental and model
210 values of $D_{\text{NiO}}^{\text{Ol/L}}$ (i.e., $(D_{\text{Ni}} \text{ XP} - D_{\text{Ni}} \text{ Model})/D_{\text{Ni}} \text{ XP}$), so that the relative root mean square
211 error is:

212

$$213 \quad \text{R RMSE} = ((D_{\text{Ni}} \text{ XP} - D_{\text{Ni}} \text{ Model})/D_{\text{Ni}} \text{ XP})^2/N)^{0.5} \quad (9)$$

214 Results are shown in Table A1 for various experimental databases. Matzen et al. (2013)
215 critically evaluated the quality of the experimental database, and recommended that data of
216 questionable quality be filtered. Using the filtered database A consisting of 325 experiments
217 with and without FeO, the Matzen et al. (2013) model has a R RMSE = 0.21, and it is more
218 accurate than the Beattie-Jones model with a R RMSE = 0.25. The experimental database

219 assembled by Herzberg et al. (2013) was unfiltered for data quality, but filtered to include data
220 with FeO as being representative of basalts, picrites and komatiites in nature; data without FeO
221 were excluded. This database (Herzberg et al. 2013) is archived in Table A2 of the Appendix,
222 and was constructed from experimental data compiled by Li and Ripley (2010) to which the data
223 of Matzen et al. (2010; 2013) were added. There are 284 experiments in Table A2, and Table A1
224 shows that the models of Matzen et al. (2013; R RMSE = 0.23) and Beattie-Jones (R RMSE =
225 0.21) are essentially of equal accuracy in describing these data. This is more easily visualized in
226 Figures 1a and 1b. Moreover, Figure 1c shows that the Matzen et al. (2013) model description of
227 the experiments is well captured by the Beattie-Jones model.
228

229 Both olivine/melt NiO partitioning models describe equally well (i.e., R RMSE = 0.21)
230 the database A of Matzen et al. (2013) that was filtered for quality and which was restricted to
231 FeO-bearing samples only (Table A1). It is not clear why FeO-free experiments yield a less
232 accurate Beattie-Jones model. However, we focus on understanding the relationship between
233 olivine Mg number (i.e., mole% 100MgO/(MgO + FeO)) and Ni content of naturally-occurring
234 olivines, and the Beattie-Jones model is clearly justified. As shown in Figure 1 and Table A1,
235 there is no intrinsic disadvantage of Beattie-Jones compared with Matzen et al. (2013) model,
236 despite the fact that it is not rooted in a thermodynamic parameterization. However, we
237 acknowledge that while the Beattie-Jones model provides an excellent description of many
238 olivine occurrences discussed below, it may fail in situations where melting and crystallization in
239 nature operate outside the experimental calibration bounds. In a section that follows, we show
240 that both nickel partitioning models do a good job of describing the Ni contents of olivines from
241 the Paleocene picrites from West Greenland and Baffin Island.
242

243 244 **NICKEL CONTENTS OF OLIVINE AND MANTLE MELTS OF A** 245 **PERIDOTITE SOURCE** 246

247 **Model Results** 248

249 The Ni contents of near-solidus primary magmas are only slightly higher than those of
250 primary magmas for L+Ol and L+Ol+Opx assemblages, and the full range of possibilities are
251 shown in Figure 2a for batch melting of fertile peridotite KR4003; however, this small difference
252 in melt Ni content is magnified by the Ni contents of coexisting olivines, which can show a range
253 of values at constant MgO and Mg-number. The compositions of olivine coexisting with liquids
254 for which olivine is the only crystalline phase (L+Ol) are shown as blue tie-lines. There is
255 considerable rotation of the tie lines through the bulk composition that illustrates mostly
256 compatible behavior of Ni in olivine. However, at MgO in the ~ 36 to 38% range the Ni content
257 of model olivine is less than the Ni content of these high MgO melts, demonstrating
258 incompatible behavior. As temperatures decrease from the liquidus to the solidus, the large drop
259 in Ni in melts is mass balanced by a large Ni gain in olivine (Figure 2a).
260

261 For accumulated fractional melting of fertile peridotite, the Ni content of a primary
262 magma is:
263

264
$$\text{Ni (ppm)} = 21.6\text{MgO} - 0.32\text{MgO}^2 + 0.051\text{MgO}^3 \quad (10)$$

265

266 for MgO contents < 27% (Herzberg 2011). The Ni contents of accumulated fractional melts of
267 fertile peridotite are slightly higher than those of batch melts at any specific MgO content, and
268 they are similar to the Ni contents of melts on the solidus (Figure 2b). Importantly, for any
269 specific mode of melting and melt MgO contents < 27%, the Ni content of a primary melt is
270 essentially identical for both fertile and depleted peridotite sources (Figure 2b).

271

272 The Ni contents of olivine in equilibrium with primary magmas of fertile peridotite are
273 given in Figure 3 at mantle pressures appropriate to partial melting, which typically range from ~
274 2 to 6 GPa for Phanerozoic magmas (Herzberg and Asimow 2015). Slight elevations in Ni for
275 accumulated fractional melts relative to batch melts (Figure 2b) are reflected in their slightly
276 higher olivine Ni contents (Figure 3a). It is important to note, however, that a computed
277 aggregate fractional melt is not in equilibrium with its residue and residual olivine; only the final
278 drop of liquid extracted is in equilibrium with the residue (Herzberg 2004). Nevertheless, it will
279 be shown below that the assumption of equilibrium is a good approximation.

280

281 Our analysis of olivines of primary magmas of a peridotite source always show an
282 maximum of Ni content with increasing melt fraction (F; Figure 3a). For the case of accumulated
283 fractional melting, the Ni content of olivine is ~ 2860 ppm at the solidus, increasing to a
284 maximum of 3100 ppm at F = ~ 0.5, and then decreasing by almost half to 1600 ppm at F = 1.
285 The large drop in olivine Ni content is nothing more than a mass balance requirement of the
286 large increase in melt Ni content from the solidus to the liquidus (Figure 2a). Our model differs
287 fundamentally from that reported by Straub et al. (2008), wherein their Ni contents increase
288 continuously with increasing degree of melting (Figure 3a). As a critical test, we compare our
289 model olivine compositions for batch melting at high pressures with olivine compositions
290 measured from experiments on peridotite KLB-1 (Zhang and Herzberg 1994; Herzberg and
291 Zhang 1996). Results are compared with that of Straub et al. (2008) and, although this model is
292 for fractional melting, it will not differ significantly from our batch melting model.

293

294 The experimental olivines are distributed between the solidus and liquidus in experiments
295 that were conducted in a temperature gradient on peridotite KLB-1, which has 0.25% NiO (see
296 Figure 4 in Zhang and Herzberg 1994). Only a few representative olivines were reported by
297 Herzberg and Zhang (1996), and we include the full dataset in Table A3 of the Appendix. There
298 is liquid migration toward the hot spot in a temperature gradient by the solution and precipitation
299 of crystalline phases dominated by olivine (e.g., Leshner and Walker 1988; Zhang and Herzberg
300 1994). Expulsion of low degree melts from near the solidus and toward the liquidus near the
301 hotspot results in mixed melts, and there is a continuum of olivine compositions in equilibrium
302 with these melts (Figure 3b). Notably, experimental olivines from melts near the liquidus have
303 the lowest Ni contents, about 1600 ppm Ni, and maximum Ni contents of about 3100 ppm
304 (Figure 3b). We have not examined the possibility that the low olivine Ni contents were
305 compromised by Ni loss to rhenium containers; however significant Ni loss is not likely because
306 there is good mass balance in Ni between near liquidus olivine and liquid compositions with
307 respect to 0.25% for the bulk KLB-1 composition, within the reported uncertainties (Herzberg
308 and Zhang 1996). This provides experimental confirmation of our model olivine compositions

309 calculated from Beattie-Jones, Toplis, and Herzberg-O'Hara parameterizations of experimental
310 data. Low Ni olivine contents near the liquidus illustrate the incompatible nature of Ni in olivine
311 with respect to near-total melts of mantle peridotite.
312

313

314 **Olivine in Mantle Peridotite**

315

316 The evidence presented in Figure 4 suggests that olivines in mantle peridotite could have
317 been in equilibrium with primary melts that formed by either batch or accumulated fractional
318 melting. This conclusion is based on the excellent match between calculated olivine
319 compositions for our assumed peridotite having 0.25% NiO (McDonough and Sun 1995) and
320 those of 215 olivine grains in mantle peridotites from a wide range of tectonic environments that
321 have been measured by high precision electron microprobe analysis (Herzberg et al. 2013).
322 These natural olivines contain 2800 to 3100 ppm Ni, and Mg numbers in the 89.5 to 92.0 range
323 (Figure 4a), and there is a positive relationship between Ni and Mg number (i.e., $\delta\text{Ni}/\delta\text{Mg}$
324 number = 68), which is captured in the computed olivine compositions. Olivines from peridotite
325 residues that have Ni compositions outside the bounds shown in Figure 4a have been reported
326 (e.g., Doucet et al. 2012), indicating the operation of second stage processes such as interaction
327 with sulfide (Barnes et al. 2013).
328

329

330 If mantle peridotite is the solidification product of an early magma ocean (Herzberg and
331 O'Hara 1985; Melosh 1990; Tonks and Melosh 1993; Walter and Trønnes 2004; Elkins-Tanton
332 2012; Carlson et al. 2014) then the evidence from olivine indicates that it melted and crystallized
333 with little or no removal of olivine (i.e., equilibrium/batch melting and crystallization; Figure
334 4b). The liquidus temperature of a totally melted fertile peridotite composition similar to
335 KR4003 at 1 atmosphere is $\sim 1700^\circ\text{C}$, and $> 2000^\circ\text{C}$ at $P > 7 \text{ GPa}$ (Herzberg and Asimow 2015).
336 The liquidus olivine to crystallize from this magma ocean composition would have an Mg
337 number of 96 and a Ni content of 1600 ppm (Figure 4b). Olivine compositions of equilibrium
338 (batch) crystallization would equilibrate to lower Mg numbers and higher Ni contents, reaching a
339 maximum of about 3000 ppm Ni after about 70% crystallization. At total solidification, the
340 fertile peridotite would have olivine containing 2860 ppm Ni and an Mg number of 89.5, similar
341 to those that have been measured (Figure 4b; Herzberg et al. 2013); olivines with higher Mg
342 numbers likely formed in a second stage as residual olivines after melt extraction (Figure 4a).

343

344 Formation of mantle peridotite by perfect removal or fractional crystallization of olivine
345 in a terrestrial magma ocean is not plausible unless there is an error in our assumed peridotite
346 composition. This process would form olivines with an array of compositions, many of which
347 would have had much higher Ni contents than observed. The red curve in Figure 4b describes
348 the olivine crystal line of descent that would result from 0 to 82% fractional crystallization of
349 olivine from a totally melted peridotite. The calculation does not include the crystallization of
350 pyroxenes, spinel and garnet, all of which would elevate the Ni contents of the magma ocean and
351 its crystallizing olivines as indicated by the red arrow because Ni is more incompatible in these
352 phases than olivine (Beattie et al. 1991; Herzberg et al. 2013); their inclusion would further
353 magnify the misfit between the olivine-only CLD and observed mantle olivines. However, the
slightly higher Ni contents of observed olivines relative to the equilibrium crystallization

354 situation (Figure 4b) might be accommodated by a small measure of fractional crystallization.
355 That is, olivine may have maintained an equilibrium composition with its magma ocean owing to
356 suspension as it crystallized, but it might not have been perfect owing to a small amount of
357 removal somewhere. The petrological evidence for the dominant operation of equilibrium
358 crystallization is consistent with convection velocities that were higher than settling velocities at
359 relatively low pressures appropriate to olivine stability (Tonks and Melosh 1990; Solomatov
360 2015). Trace element and isotopic evidence indicate that fractional crystallization might have
361 been more important in the lower mantle, involving perovskite and ferropericlasite (Caro et al.
362 2005; Corgne et al. 2005; Labrosse et al. 2007; Rizzo et al. 2011; Puchtel et al. 2013). This might
363 be manifest by lithological and geochemical heterogeneity at D'' (Labrosse et al. 2007; Coltice et
364 al. 2011), in contrast with the remainder of the mantle from which peridotite is sampled.
365

366 A corollary to the formation of Earth's mantle by equilibrium-dominated crystallization
367 is that its magma ocean solidified to a peridotite mantle its own composition, with little or no
368 basaltic type differentiates and dunite cumulates anywhere. This implies the Earth's surface was
369 largely made of peridotite at some early stage, a conjecture that might be consistent with some
370 magma ocean models that suggest the Earth cooled from a total melt to a complete solid in a
371 short period of time, that is 10^6 - 10^7 years (Hamano et al. 2013; Elkins-Tanton 2008). However,
372 these estimates are highly uncertain because there is currently no comprehensive physical model
373 of magma ocean crystallization that considers the time scales of crystal nucleation, growth and
374 separation from liquid (Solomatov 2015). Nevertheless, the evidence from the Ni content of
375 olivine in mantle peridotite points to a major role for equilibrium crystallization in an early
376 turbulently convecting magma ocean; subsequent stages of partial melting operated to slightly
377 elevate Ni and Mg number in residual olivines. Uncertainties in the depths and temperatures of
378 magma ocean melting and crystallization provide a good example of the utility of using the
379 Beattie-Jones model for understanding the Ni content of olivine.
380

381 382 **Olivine in Archean Komatiites**

383
384 Archean-age komatiites from Alexo and Munro in the Abitibi greenstone belt have
385 compositions that indicate extreme degrees of melting of a peridotite source. Petrological
386 analysis indicates all crystalline phases were melted out except olivine (Herzberg 2004), and
387 primary magma compositions have been estimated to have contained 26-30% MgO (Arndt et al.
388 2008; Herzberg 2011). The major element geochemistry of Alexo and Munro komatiites is
389 consistent with accumulated fractional melting (see Figure 11 in Herzberg 2004).
390

391 Olivine compositions from Alexo and Munro komatiites are shown in Figure 5 together
392 with those that have been modeled for olivines that crystallize at the surface from primary
393 accumulated fractional melts of a peridotite source (Herzberg 2011). As shown more clearly
394 below and in Figure 6, model olivines of primary magmas that crystallize at the surface have
395 slightly higher Mg numbers than do those that crystallize in the mantle. Many olivines from
396 Alexo contain about 3300 ppm Ni and Mg numbers of 94.5 (Sobolev et al. 2007), in good
397 agreement with those expected to crystallize from a primary magma having 30% MgO; other
398 olivines are more consistent with 26% MgO. In some cases, observed olivines have Ni contents

399 that are up to ~ 500 ppm higher than those expected of primary magmas, and this might indicate
400 a pyroxenite contribution to melting (Sobolev et al. 2007). However, the pyroxenite
401 interpretation is not consistent with olivine Mn and Ca contents (Figure 7b in Herzberg 2011),
402 indicating that the komatiites melted from a peridotite source that was free of recycled crust.
403 Instead, Alexo and Munro olivines have Ni contents that are similar to model olivines produced
404 by partial melting of a normal peridotite source having 1964 ppm Ni, followed by fractional
405 crystallization of olivine. Importantly, calculated olivine crystal lines of descent capture the
406 curvature to the Ni-Mg number systematics in measured olivines (Figure 5; Herzberg 2011). For
407 example, fractional crystallization of olivine from a primary magma having 30% MgO would
408 drive up the Ni content of olivine from 3040 to 3660 ppm (Figure 5). This unusual behavior
409 contrasts with the more familiar drop in Ni along an olivine CLD as seen in Figure 6 for MORB
410 primary magmas with lower MgO contents.

411
412 The volatile-free liquidus temperature of a primary magma having 30% MgO is 1600°C
413 at 1 atmosphere, and > 1800°C at $P > 5$ GPa (Herzberg and Asimow 2015). Based on the phase
414 diagram for KLB-1 (Herzberg and Zhang 1996), melting might have commenced in the lower
415 mantle or transition zone, and may be consistent with mantle potential temperatures in the 1800
416 to 2000°C range (Herzberg et al. 2010; Putirka 2016). Archean komatiites have higher MgO
417 contents than the more common contemporaneous basalts that melted from ambient mantle
418 (Herzberg et al. 2010), consistent with a hot plume origin (e.g., Arndt et al. 2008; Herzberg et al.
419 2010). However, this conclusion assumes that volatile-free melting is valid. Evidence has been
420 presented for a role played by CO₂ and H₂O in melting that produced the komatiites from Alexo
421 and Pyke Hill (Sobolev et al. 2016), and there is ambiguity concerning the mantle potential
422 temperature of the source (Herzberg 2015). Melting of peridotite that was carbonated will be
423 extensive and highly productive owing to the effect of CO₂ in lowering the solidus temperature
424 (Dasgupta 2013); unfortunately, the extent to which degassing has affected CO₂ and H₂O is not
425 known (Herzberg 2015), and the depths of melting may have been located anywhere from the
426 lower to the upper mantle. Uncertainties in the depths and temperatures of melting provide
427 another good example of the advantage of using the Beattie-Jones model for understanding the
428 nickel content of olivine.

429
430

431 **Olivine from Mid-Ocean Ridge Basalts**

432

433 New high precision electron microprobe data have been obtained on olivines from the
434 Siqueiros fracture zone at the East Pacific Rise (Table A4) using analytical methods described in
435 the Appendix. Results are shown in Figure 6 together with high precision data from the
436 Southeast Indian Rise (Sobolev et al. 2007). The trend of olivine analyses matches reasonably
437 well the computed CLD for a representative MORB primary melt composition from the
438 Siqueiros fracture zone having 11.5% MgO (Herzberg and Asimow 2015). The most primitive
439 olivines have Ni contents that trend towards 2900 ppm, similar to those of model olivines from
440 primary magmas. Those olivines displaying low Ni contents relative to the CLD might indicate
441 sequestration from a sulfide phase (Herzberg 2011); high Ni olivines might be attributed to
442 magma chamber crystallization, recharge, mixing, and tapping (O'Hara 1977; Coogan and
443 O'Hara 2002; O'Hara and Herzberg 2002; O'Neill and Jenner 2012). However, the overall good

444 agreement between observed and calculated olivine Ni contents indicates a robust computational
445 method constructed from Beattie-Jones, Toplis, and Herzberg-O'Hara partitioning
446 parameterizations. And the success of this model to mantle peridotite (Figure 4b; Ol liquidus T
447 at 1 atmosphere = 1700°C), Archean komatiites (Figure 5), and present-day MORB (Figure 6; Ol
448 liquidus T at 1 atmosphere = 1280°C) is a good illustration of its general applicability to an
449 extremely wide range of temperature and pressure environments on Earth.

450
451

452 **Olivine from Iceland: The Modern Icelandic Mantle Plume**

453

454 Iceland is the modern expression of volcanism from a mantle plume that can be tracked
455 back in time to its first eruption in the Paleocene (Saunders et al. 1997). Olivine compositions
456 from Theistareykir volcano are shown in Figure 7, and they are representative of the full range of
457 possibilities from other volcanoes on Iceland (Sobolev et al. 2007). There are two major
458 populations of data. The first can be reasonably described by model olivines that crystallized for
459 the equilibria L+Ol and L+Ol+Plag (Figure 7). The second olivine population has a lower Mg
460 number at constant Ni content, and is roughly coincident with the curved red lines shown in
461 Figure 7. These red lines represent olivines that crystallize from magmas that form by the
462 mixing of the primary magma with derivative magmas along the LLD. This is a simulation of
463 magma chamber crystallization, recharge, and mixing (O'Hara 1977) and it shares some
464 similarities to those for MORB (O'Neill and Jenner 2012; Coogan and O'Hara 2015), Mangaia
465 and Curacao picrites (Herzberg et al. 2014; Trela et al. 2015). Although the calculations were
466 performed at 1 atmosphere, the effect of increasing the pressure to those expected in the deeper
467 crust is to shift the Ni contents of olivines on the L+Ol+Plag+Cpx CLD, but this will not affect
468 the mixing lines.

469

470 Figure 7 illustrates the very good agreement between observed and calculated olivine Ni
471 contents using Beattie-Jones, Toplis, and Herzberg-O'Hara parameterizations. White circles are
472 olivine compositions that have been calculated using the method of Matzen et al. (2013) for
473 $D_{\text{NiO}}^{\text{Ol/L}}$ and with the assumption of 2975 ppm Ni in the peridotite source olivine (Figure 4).
474 Matzen et al. (2013; 2016) have suggested that the Ni contents of olivine phenocrysts will be
475 elevated for the condition of higher temperatures and pressures of magma generation in the
476 mantle compared with the temperature of crystallization in the crust, a ΔT effect. This is shown
477 in Figure 7 for melting pressures that range from 0 to 2.1 GPa. Indeed, the Matzen model is
478 sensitive to assumptions concerning the conditions of melting. Flipping this around, there is
479 great potential for using the Matzen method for independently estimating final pressures of melt
480 segregation in the mantle. However, it is only for a melt segregation pressure of 0 GPa where
481 there is excellent agreement with both observed and calculated olivines using the Beattie-Jones
482 parameterization, an implausible result because melting must be deeper than the Icelandic crust,
483 which is 25 to 40 km thick (Darbyshire et al., 2000). A melting pressure of 1.0 GPa yields a
484 higher Matzen Ni content for olivine, but it is well within the uncertainty of the partitioning
485 models. However, these Matzen pressures depend on the assumed Ni content of the olivine in
486 the residual mantle. If we drop Ni from 2975 Ni to 2710 ppm, then a good match can be obtained
487 with observed olivines at an assumed a pressure of 1.0 GPa below Iceland. Olivine in mantle

488 peridotite with 2710 ppm Ni is lower than the average of our high precision data and the
489 expectation of 1964 ppm Ni in fertile peridotite (Figure 4), but it is a reasonable low Ni bound.
490

491 There are two important messages to be drawn from Figure 7. The first is that olivines
492 from Iceland have Ni contents that are well described by model olivines expected by melting of a
493 normal peridotite source having a canonical Ni content of 1964 ppm (McDonough and Sun
494 1995). Icelandic olivines do not have Ni contents in excess of 3000 ppm, differing from those
495 observed from Hawaii and for which pyroxenite source melting has been inferred (e.g., Figure 7;
496 Sobolev et al. 2007; Herzberg 2011); a peridotite source for Iceland inferred from Ni is also
497 consistent with high Ca and Mn, and low Fe/Mn, in contrast with Hawaii (please see below;
498 Herzberg et al. 2013). The second message is that inferences about source lithology can be
499 compromised without due consideration given to magma chamber crystallization, recharge, and
500 mixing. For example, the Theistareykir olivine population with low Mg number might be
501 erroneously interpreted as requiring the involvement of a modest amount of pyroxenite as it is
502 shifted in the direction towards Hawaii; but to reiterate, this is not consistent with a peridotite
503 source as expressed by olivine phenocrysts having high Ca and Mn, and low Fe/Mn (see below).
504

505 We conclude that the evidence from olivine chemistry points to little or no role is played
506 by recycled crust as a distinct lithology in the present-day Icelandic mantle plume. In contrast,
507 the enriched and depleted geochemical variability (e.g., Fitton et al. 2003) was interpreted by
508 Shorttle et al. (2014) as arising from the melting a lithologically heterogeneous source consisting
509 of recycled crust imbedded in buoyant harzburgite. We suggest herein that another possibility
510 might be a source that consists of harzburgite with variable amounts of recycled crust mixed in.
511 It was noted previously that the Ni contents of primary melts (i.e., MgO < 27%) of peridotite
512 having 38 to 42% MgO are essentially identical (Figure 2b), as will be the Ni contents of their
513 crystallizing olivines. Therefore, olivine phenocryst chemistry cannot be used to distinguish a
514 fertile peridotite source from a depleted harzburgitic source that had been variably refertilized by
515 mixing moderate amounts of recycled crust.
516

517 Mantle plumes may be anchored at the edges of Large Low Shear Velocity Provinces on
518 the core-mantle boundary (Burke et al. 2008), and LLSVPs may consist of ambient mantle mixed
519 with recycled crust (Mulyukova et al. 2015), primordial mantle differentiates (Labrosse et al.
520 2007), or both. Recycled crust from the LLSVP was sampled by the Hawaiian mantle plume
521 (Weis et al. 2011) which stretched it into filaments (Farnetani and Hofmann 2009), but it was not
522 mixed or destroyed as a distinct pyroxenite lithology as indicated by high Ni olivine in Figure 7
523 (Sobolev et al. 2007; Herzberg 2011). From a petrological point of view, Iceland does not
524 conform to this standard model as its olivine Ni contents are low and consistent with peridotite
525 melting (Figure 7). Mangaia is another example, although in detail the olivine compositions
526 differ from those of Iceland (Herzberg et al. 2014). The question is how exactly does recycled
527 crust get mixed into its peridotite host? Destruction of recycled crust may occur by partial or
528 total melting of the pyroxenite and injection of its silicic melts into the surrounding mantle
529 peridotite, yielding a relatively refertilized peridotite (Yaxley and Green 1998). Such a source
530 can have phantom-like properties in that its lithological identity as recycled crust can be
531 destroyed while its trace element and isotope ratios are preserved in the refertilized peridotite
532 (Herzberg et al. 2014). Whether pyroxenite is destroyed as in Iceland or preserved as in Hawaii

533 will depend on a potentially complex interplay amongst its mass relative to the peridotite host,
534 the thermal properties of the mantle plume, and the thickness of the lithosphere. The signal of
535 pyroxenite melting in olivine chemistry is likely to be stronger for mantle plumes that impact a
536 thick lithosphere (Sobolev et al. 2007) because pyroxenite has a lower solidus temperature and
537 higher melt productivity relative to peridotite (e.g., Morgan 2001; Pertermann and Hirschmann
538 2003; Ito and Mahoney 2005; Sobolev et al. 2007; Shorttle and MacLennan 2011). The present-
539 day Icelandic mantle plume is only slightly cooler than Hawaiian plume (Herzberg and Asimow
540 2015), but it is coincident with the Mid Atlantic ridge; therefore, the Icelandic melting column is
541 drawn closer to the surface relative to Hawaii, and there are increased opportunities for
542 pyroxenite destruction and peridotite refertilization.

543
544

545 **Olivine from the Paleocene Picrites of West Greenland (WG) and Baffin Island (BI): The** 546 **Ancient Icelandic Mantle Plume**

547

548 Paleocene picrites from West Greenland and Baffin Island were the first magmatic
549 eruptions from the ancestral Icelandic mantle plume (Saunders et al. 1997), and they are notable
550 in having the most primitive He, Pb and Nd isotopic compositions on Earth (Stuart et al. 2003;
551 Starkey et al. 2009; Jackson et al. 2010). Inversion of the major elements indicate that they
552 crystallized from high MgO lavas that were generated by a hot mantle plume source (Larsen and
553 Pedersen 2000; Herzberg and O'Hara 2002; Herzberg and Gazel 2009; Hole and Millet 2016).
554 The implication is that the ancestral Icelandic mantle plume sampled a primitive region of
555 Earth's mantle.

556

557 Low and high precision data (Larsen and Pedersen 2000; Sobolev et al. 2007) show that
558 olivines from WGBI lavas have Ni contents that are higher than those expected of melting of a
559 peridotite source having a canonical Ni content of 1964 ppm (Herzberg et al. 2013). This result
560 has been reproduced and illustrated again in Figure 8 with new high precision electron
561 microprobe data that we report in Table A5 of the Appendix. The trend of WGBI olivines is not
562 consistent with the expectations of a crystal line of descent involving L+Ol and L+Ol+Plag.
563 Crystallization of clinopyroxene is indicated, followed by mixing with more primitive melts as
564 for Theistareykir volcano. Importantly, the signal of clinopyroxene crystallization will be
565 revealed in olivines with elevated Ni, but the rocks may contain no clinopyroxene as a
566 phenocryst phase. Together with MORB (Figure 6), Mangaia (Herzberg et al. 2014), Curacao
567 (Trela et al. 2015), and Theistareykir (Figure 7), olivine from West Greenland and Baffin Island
568 is faithfully recording in its chemistry the widespread occurrence of magma chamber fractional
569 crystallization, recharge, and mixing (O'Hara 1977; Coogan and O'Hara 2015; O'Hara and
570 Herzberg 2002; O'Neill and Jenner 2012).

571

572 Some olivines from Baffin Island with Mg numbers in the 85-88 range have Ni contents
573 that are similar to olivines that crystallized from mixed magmas of Theistareykir volcano;
574 however, the same BI olivine population with Mg numbers > 90 is always higher in Ni by as
575 much as 500 ppm in some cases, indicating a fundamental difference with respect to the modern
576 Icelandic mantle plume. One possible explanation is that Ni is high in WGBI olivines because of
577 the participation of recycled crust. However, while recycled crust is likely to be compositionally

578 heterogeneous (Herzberg 2011), olivines that crystallize from melts of a pyroxenite source
579 commonly have low Ca and Mn contents, and elevated Fe/Mn; this signature can arise from
580 retention of Ca and Mn in modally abundant residual clinopyroxene and garnet, respectively
581 (Herzberg 2011). This is in contrast with olivines from WGBI that have Mn, Fe/Mn, and Ca
582 contents similar to those of olivines in MORB (Herzberg et al. 2013), as is shown in Figure 9.
583 Fe/Mn for olivines from WGBI and Theistareyker volcano are similar to those for MORB and
584 model olivines expected of a peridotite source; they are also much lower than Fe/Mn of olivines
585 from Hawaii for which a pyroxenite source has been interpreted (Sobolev et al. 2007; Herzberg
586 2011).

587
588 In general, olivines from West Greenland have slightly higher Ni contents than do those
589 from Baffin Island (Figure 8). No correlation exists between excess Ni in olivine and the
590 isotopic compositions of the lavas from which our olivine analyses were obtained (Starkey et al.
591 2009). The only correlation observed in trace element ratios is that West Greenland lavas
592 generally have higher V/Sc than do Baffin Island lavas (Figure 10). This is based on histograms
593 of V/Sc for West Greenland and Baffin island lavas (Larsen and Pedersen 2009; Lightfoot et al.
594 1997; Starkey et al. 2009) that have been filtered with Nd and Sr isotope ratios to exclude
595 samples that had experienced crustal contamination (Larsen and Pedersen 2009). Increasing fO_2
596 increases V/Sc in melts of spinel and garnet peridotite (Lee et al. 2005), but the effect is greater
597 for garnet peridotite (Mallmann and O'Neill 2009) at constant fO_2 . This is consistent with
598 petrological modeling that shows overlapping but generally greater depths of melting in the
599 generation of West Greenland primary magmas compared with Baffin Island (Hole and Millett
600 2016).

601
602 The Beattie-Jones Ni partitioning model for a McDonough and Sun (1995) peridotite
603 predicts olivines with lower Ni contents than those observed for WGBI, but the misfit is within
604 the uncertainty of the model. Matzen et al. (2013; 2016) have suggested that Ni contents of
605 olivine phenocrysts that crystallize at the surface will be elevated for hotter melts generated
606 deeper in the mantle, the ΔT effect. This interpretation is consistent with higher Ni contents of
607 WG olivines compared with BI olivines (Figure 8). Olivines represented by the white circles in
608 Figure 8 have been calculated using the method of Matzen et al. (2013) for $D_{NiO}^{Ol/L}$ assuming an
609 olivine Ni content of 2975 ppm Ni in the peridotite source (Figure 4), melting pressures that
610 range from 2.8 to 2.1 GPa, and crystallization at the surface (see caption for details). The
611 agreement with many WGBI olivines is very good for the high Ni population (Figure 8). Some
612 of the low Ni BI olivine population is consistent with a Matzen melting pressure of 2.1 GPa;
613 however, BI olivines with the lowest Ni contents would likely require unreasonably low
614 pressures or Ni-depleted peridotite.

615
616 How do we explain Icelandic and WGBI olivines that have similar Fe/Mn (Figure 9) but
617 different Ni contents (Figures 7, 8)? This paradox was explained by Paleocene picrites that
618 melted from a nickel-rich peridotite source (Figure 8), Ni being high owing to interaction with
619 the core (Herzberg et al. 2013). However, this interpretation is no longer unique because the
620 Beattie-Jones $D_{NiO}^{Ol/L}$ model was assumed. In contrast, the nickel partitioning model of Matzen
621 et al. (2013) is consistent with elevated Ni in olivine owing to deep and hot melting of normal
622 mantle peridotite having the canonical 0.25% NiO (McDonough and Sun 1995) and containing

623 residual olivines with 2975 ppm Ni (Figures 4). The nickel-rich peridotite hypothesis (Herzberg
624 et al., 2013) is now less secure, but Ni contents of olivine using the Matzen et al. (2013) method
625 are sensitive to assumed T-P conditions of melting and Ni content of residual olivine. We
626 caution, however, against over interpretation because the uncertainties in both Beattie-Jones and
627 Matzen models are about the same (Figure 1; Table A1), and they propagate to significant
628 uncertainties in the Ni content of olivine as shown in Figure 8. More work is needed to constrain
629 the problem.

630
631

632 **Olivine from Hawaii**

633

634 Picrites from the West Greenland Vaigat Formation were proposed to have segregated
635 below a thick 100-120 km continental lithosphere (Larsen and Pedersen 2000), consistent with
636 model pressures reported by Hole and Millet (2016). These pressures are comparable to those
637 below Hawaii (Li et al. 2000; 2004), and mantle potential temperatures for WG and Hawaii are
638 are about the same ($T_p = 1500\text{-}1600^\circ\text{C}$; Herzberg and Gazel 2009; Hole and Millet 2016).
639 Therefore, it is expected that the model of Matzen et al. (2013; 2016) would predict olivine Ni
640 contents from West Greenland picrites that are high like those from Hawaii owing to a
641 comparable ΔT effect; however, this prediction is not consistent with measured Ni contents
642 (Figure 8). One possibility is that Hawaii melted from a peridotite source that was unusually rich
643 in nickel. However, this interpretation is not unique because the higher Ni contents of Hawaiian
644 olivines, together with lower Ca and higher Fe/Mn, are also consistent with the expectations of
645 pyroxenite melting (Sobolev et al. 2005; 2007; Herzberg 2011). Furthermore, both olivines and
646 whole rock lava compositions in Mg-rich shield stage Hawaiian lavas are low in CaO; estimated
647 high MgO primary melt compositions have lower CaO contents than experimental and
648 thermodynamic melts of mantle peridotite (Walter, 1998; Herzberg 2006; Herzberg and Asimow
649 2008; Jennings and Holland 2015), consistent with pyroxenite melting.

650

651

652 **COBALT CONTENTS OF OLIVINE AND MANTLE MELTS OF A** 653 **PERIDOTITE SOURCE**

654

655 **Model Results**

656

657 Compared with nickel, cobalt data for olivine are lacking in quality and quantity owing to
658 its lower abundance levels. Also, there has never been a comparable understanding of how
659 cobalt is partitioned between melts and olivine. Building on the success of the nickel modeling
660 as discussed above, we will finish this contribution with a preliminary cobalt model.

661

662 Cobalt is more incompatible in olivine than Ni, but the partition coefficients depend on
663 composition (equations 4 and 5). Cobalt contents of primary and derivative magmas and their
664 crystallizing olivines have been calculated using the same procedure discussed above for nickel,
665 using the partition coefficient from equation (5) (Beattie et al. 1991). These are compared with
666 new high precision electron microprobe analyses of San Carlos olivine we provide in Table A6;
667 our analytical method is discussed in the Appendix. We have found small differences in two San

668 Carlos olivine grain populations; one grain has an Mg number of 90.03 (± 0.06 ; 1σ) and 139 (± 7 ;
669 1σ) ppm Co; another has an Mg number of 90.38 (0.04; 1σ) and 142 ppm (± 7 ; 1σ) Co. These
670 data are similar to those reported by Sobolev et al. (2007; Mg number 90.22 and 140 ppm Co),
671 and to 142 ppm Co reported by De Hoog et al. (2010) from LA-ICP-MS. Taken at face value,
672 these data suggest a slight positive correlation of Mg number and Co content. Regression yields
673 a model olivine at the solidus with an Mg number of 89.5 and 135 ppm Co, and we use this in
674 our modeling. We assume peridotite KR4003 has a 105 ppm Co (McDonough and Sun 1995).
675

676 Primary magma solutions to the equation for accumulated fractional melting are shown in
677 Figure 11a. Liquids extracted from dunite [L + Ol] and harzburgite [L + Ol + Opx] residues have
678 similar Co contents at any specific MgO content. Co contents of liquids on the solidus can be
679 higher by ~ 10 ppm for high MgO melts; Co contents of liquids extracted from lherzolite are an
680 interpolation. The compositions of olivines in equilibrium with these primary magmas of
681 accumulated fractional melting are shown in Figure 11b.
682

683 **Olivine from Mantle Peridotite**

684
685
686 Computed olivine compositions in equilibrium with primary magmas produced by
687 accumulated fractional melting at mantle pressures are shown in Figure 12. These are in
688 excellent agreement with those that have been measured for San Carlos olivine reported by
689 Sobolev et al. (2007) and our new high precision electron microprobe data (Table A6). The
690 computed olivine compositions are also in good agreement with olivine compositions from
691 orogenic peridotite and xenoliths that have been measured by LA-ICP-MS (De Hoog et al.
692 2010). As with Ni, the agreement between observed and calculated olivine Co contents indicates
693 a robust computational method constructed from Beattie-Jones, Toplis, and Herzberg-O'Hara
694 parameterizations of experimental data.
695

696 **Olivine from the Siqueiros Transform of the East Pacific Rise**

697
698
699 High precision electron microprobe Co measurements made for olivines from the
700 Siqueiros Transform are given in Table A4 and shown in Figure 13 together with those from the
701 Knipovich ridge (Sobolev et al., 2007). Siqueiros olivine Co contents are in good agreement
702 with those reported by Sobolev et al. (2007), and they are consistent with ~ 140 ppm Co
703 calculated for olivines in equilibrium with primary magmas having 10-13% MgO. However, the
704 olivine-only CLD at 1 atmosphere has generally lower Co than those which have been reported
705 for olivines from the Knipovich MORB (Figure 13b; Sobolev et al. 2007). Similarly the
706 calculated olivine-only LLD is lower in Co than those which have been measured for MORB
707 glasses (Figure 13a; Jenner and O'Neill, 2012). We expect the elevated Co in both glasses and
708 olivines are the result of plagioclase and clinopyroxene fractionation, but there are no
709 partitioning data that permit a forward model and test of this conjecture.
710

711
712

713 **Olivine from the Paleocene Picrites of West Greenland and Baffin Island (WGBI): The**
714 **Ancient Icelandic Mantle Plume**

715

716 New high precision Co analyses (Table A5) for olivines from West Greenland and Baffin
717 Island lavas are compared with model olivines that crystallize from primary magmas
718 (accumulated fractional melting) at 1 atmosphere. The primitive WGBI olivines with Mg
719 numbers of 90.0 – 92.5 have Co contents that differ by about ± 10 ppm from the model olivines
720 expected to crystallize from primary magmas (Figure 14). Olivines in some samples display
721 coherent and clearly separable crystal line of descent trends, such as DUR3, DUR6 and PAD8,
722 which are larger than the ± 7 ppm (1σ) uncertainty in the precision of our San Carlos olivines. In
723 other samples, the scatter in Co is similar to the uncertainty in precision. Notably, the slightly
724 higher Ni contents for West Greenland olivines compared with those from Baffin Island are not
725 evident in Co.

726

727 Whole rock Co and MgO data reported by Larsen and Pedersen (2009) display a curved
728 trend that coincides with a computed olivine-only LLD that extends from a primary magma
729 having 20% MgO (Figure 14a). This MgO content is similar to the range of primary magma
730 composition inferred from PRIMELT2 (Herzberg and Gazel 2009) and PRIMELT3 modeling
731 (Hole and Millett 2016), which require a hot mantle plume source.

732

733 It is possible that the apparently well-resolved olivine Co contents for DUR3, DUR6 and
734 PAD8 contain information about origin. In particular, an independent evaluation is needed of the
735 model that WGBI lavas melted from a nickel-rich peridotite source by interaction with the core
736 (Herzberg et al. 2013), as discussed above. This model predicts that the peridotite source might
737 also be enriched in Co compared with the canonical 105 ppm because the partitioning of Co, like
738 Ni, between silicate and sulfide liquids is pressure dependent (Siebert et al. 2012). However, this
739 conclusion is based on a parameterization of experimental data up to 70 GPa, and it may not be
740 extrapolated with confidence to 135 GPa at the core-mantle boundary. Furthermore, the effects
741 of temperature, pressure, oxygen fugacity, and composition can add to significant uncertainties
742 in the distribution coefficients for Ni and Co between silicate and metal liquid (Walter and
743 Cottrell 2013). The main conclusion is that the rough agreement between observed and
744 calculated olivine Co is supportive of the Beattie-Jones, Toplis, and Herzberg-O'Hara
745 parameterizations to within ± 20 ppm (1σ).

746

747

748

IMPLICATIONS

749 Forward models of partial melting and crystallization of a peridotite source using existing
750 olivine-liquid partitioning for Ni, Co, Fe, Mn, and Mg (Beattie et al. 1991; Toplis 2005; Herzberg
751 and O'Hara 2002; Matzen et al. 2013) constrain the origins of measured olivines from mantle
752 peridotite and a wide range of related igneous rocks. Cobalt contents of olivine are significantly
753 lower than are Ni contents, and there are fewer high precision Co data. Implications about origin
754 are based mostly on high precision electron microprobe analysis of Ni, Fe, Mn, and Mg, and
755 there is generally good agreement where high precision Co data are available.

756

757 Olivines in mantle peridotite have Ni contents and Mg numbers that were largely
758 determined by an initial stage of equilibrium crystallization in a turbulently convecting early
759 Earth magma ocean; subsequent stages of partial melting operated to slightly elevate Ni and Mg
760 number in peridotite residues and its olivines.

761

762 Olivines from Archean komatiites from the Abitibi greenstone belt have Ni contents and
763 Mg numbers that are consistent with extensive melting peridotite source at great depths in the
764 mantle. This olivine chemistry is also consistent with primary magmas that contained 30%
765 MgO.

766

767 Some olivines from mid-ocean ridge basalts have Ni and Co contents and Mg numbers
768 that are similar to those of model olivines from primary magmas having ~ 11.5% MgO;
769 however, most have compositions that can be attributed to magma chamber crystallization,
770 recharge, and mixing (Coogan and O'Hara 2002; O'Hara and Herzberg 2002; O'Neill and
771 Jenner), in addition to sequestration from a sulfide phase (Herzberg 2011).

772

773 Olivines from the present-day Icelandic mantle plume have compositions that are
774 consistent with magma chamber crystallization, recharge, and mixing imposed on a primary
775 magma having about 17.6% MgO that formed by melting a peridotite source; unlike Hawaii, the
776 melting of recycled crust as a distinct pyroxenite lithology is not evident in olivine chemistry of
777 Iceland.

778

779 Paleocene picrites from Baffin Island and West Greenland from the ancient Icelandic
780 plume have olivines with Ni contents that are consistent with either Ni-rich peridotite that
781 formed by core-mantle interaction (Herzberg et al. 2013) or by low pressure crystallization of hot
782 and deep magmas (Matzen et al. 2013).

783

784 Olivine faithfully records in its chemistry the widespread occurrence of magma chamber
785 fractional crystallization, recharge, and mixing (O'Hara 1977). These processes typically elevate
786 the Ni content of olivine, and can potentially compromise interpretations of source lithology.

787

788 In summary, the partitioning of Ni, Mg and Fe between olivine and liquid provides a
789 successful foundation for understanding the compositions of naturally-occurring olivine that
790 formed in an extremely wide range of temperature and pressure conditions on Earth. In general,
791 hot magma oceans, mantle plumes, and ambient mantle magmatism form in ways that are
792 captured by the compositions of the olivine crystals that they contained.

793

794

795

796 ACKNOWLEDGMENTS

797

798 We are honored and grateful to Keith Putirka for his kind invitation to contribute this paper to
799 the Centennial Issue of the American Mineralogist, and for a thoughtful review. Mike Garcia is
800 thanked for editorial handling. Fin Stuart is thanked for collecting the Baffin Island samples from
which thin sections were made, and Lotte Larsen is thanked for comments. Michael Perfit is

801 thanked for providing samples from the Siqueiros fracture zone, and Michael Baker is thanked
802 for comments. We are most grateful to Andrew Matzen and Chusi Li for critical reviews.

803

804

805

REFERENCES CITED

806

807 Ahern, J.L., and Turcotte, D. (1979) Magma migration beneath an ocean ridge. *Earth and*
808 *Planetary Science Letters*, 46, 115-122.

809

810 Arndt, N.T., Leshar, C.M., and Barnes, S.J. (2008) *Komatiite*. Cambridge University Press,
811 Cambridge, 467 pp.

812

813 Barnes, S.J., Godel, B., Güreş, D., Brenan, J.M., Robertson, J., and Paterson, D. (2013) Sulfide-
814 olivine Fe-Ni exchange and the origin of anomalously Ni rich magmatic sulfides. *Economic*
815 *Geology*, 108, 1971-1982.

816

817 Beattie, P., Ford, C., and Russell, D. (1991) Partition coefficients for olivine-melt and
818 orthopyroxene-melt systems. *Contributions to Mineralogy and Petrology*, 109, 212-224.

819 Burke, K., Steinberger, B., Torsvik, T.H., and Smethurst, M.A. (2008) Plume generation zones at
820 the margins of Large Low Shear Velocity Provinces on the core-mantle boundary. *Earth and*
821 *Planetary Science Letters*, 265, 49-60.

822 Carlson, R.W., Garnero, E., Harrison, T.M., Li, J., Manga, M., McDonough, W.F.,
823 Mukhopadhyay, S., Romanowicz, B., Rubie, D., Williams, Q., and Zhong, S. (2014) How did
824 early Earth become our modern world? *Annual Review of Earth and Planetary Sciences*, 42, 151-
825 178.

826 Caro, G., Bourdon, B., Wood, B.J., Corgne, A. (2005) Trace-element fractionation in Hadean
827 mantle generated by melt segregation from a magma ocean. *Nature*, 436, 246-249.

828 Coogan, L.A., and O'Hara, M.J. (2015) MORB differentiation: *In situ* crystallization in
829 replenished-tapped magma chambers. *Geochimica et Cosmochimica Acta*, 158, 147-161.

830

831 Coltice, N., Moreira, M., Hernlund, J., and Labrosse, S. (2011) Crystallization of a basal magma
832 ocean recorded by Helium and Neon. *Earth and Planetary Science Letters*, 308, 193-199.

833 Corgne, A., Lieske, C., Wood, B.J., Rubie, D.C., and Frost, D.J. (2005) Silicate perovskite-melt
834 partitioning of trace elements and geochemical signature of a deep perovskite reservoir.
835 *Geochimica et Cosmochimica Acta*, 69, 485-496.

836

837 Darbyshire, F.A., White, R.S., and Priestley, K.F. (2000) Structure of the crust and uppermost
838 mantle of Iceland from a combined seismic and gravity study. *Earth and Planetary Science*
839 *Letters* 181, 409-428.

840

- 841 De Hoog, J.C.M., Gall, L., and Cornell, D.H. (2010) Trace-element geochemistry of mantle
842 olivine and application to mantle petrogenesis and geothermobarometry. *Chemical Geology*,
843 270, 196-215.
844
- 845 Doucet, L.S., D.A., Ionov, A.V. Golovin, and N.P. Pokhilenko (2012) Depth, degrees and
846 tectonic settings of mantle melting during craton formation: inferences from major and trace
847 element compositions of spinel harzburgite xenoliths from the Udachnaya kimberlite, central
848 Siberia. *Earth and Planetary Science Letters*, 359-360, 206-218.
849
- 850 Elkins-Tanton, L.T. (2008) Linked magma ocean solidification and atmospheric growth for Earth
851 and Mars. *Earth and Planetary Science Letters*, 271, 181-191.
852
- 853 Farnetani, C.G., and Hofmann, A.W. (2009) Dynamics and internal structure of a lower mantle
854 plume conduit. *Earth and Planetary Science Letters*, 282, 314-322.
855
- 856 Filiberto, J., Jackson, C., Le, L., and Treiman, A.H. (2009) Partitioning of Ni between olivine
857 and an iron-rich basalt: experiments, partition models, and planetary implications. *American*
858 *Mineralogist*, 94, 256-261.
- 859 Fitton, J.G., Saunders, A.W., Kempton, P.D., and Hardarson, B.S. (2003) Does depleted mantle
860 form an intrinsic part of the Icelandic plume? *Geochemistry, Geophysics, Geosystems* 4, 1032,
861 doi:10.1029/2002GC000424.
- 862 Hamano, K., Abe, Y. and Genda, H. (2013) Emergence of two types of terrestrial planet on
863 solidification of magma ocean. *Nature*, 497, 607-610.
864
- 865 Hart S.R., and Davis ,K.E .(1978) Nickel partitioning between olivine and silicate melt. *Earth*
866 *and Planetary Science Letters*, 40, 203–219.
867
- 868 Herd, C.D.K., Dwarzki, R.E.D., and Shearer, C.K. (2009) The behavior of Co and Ni in olivine
869 in planetary basalts: an experimental investigation. *American Mineralogist*, 94, 244-255.
870
- 871 Herzberg, C. (2004) Geodynamic information in peridotite petrology. *Journal of Petrology*, 45,
872 2507-2530.
873
- 874 Herzberg, C. (2006) Petrology and thermal structure of the Hawaiian plume from Mauna Kea
875 volcano. *Nature*, 444, 605-609.
876
- 877 Herzberg, C. (2011) Identification of Source Lithology in the Hawaiian and Canary Islands:
878 Implications for Origins. *Journal of Petrology*, 52, 113-146.
879
- 880 Herzberg, C. (2015) Petrological evidence from komatiites for an early Earth carbon and water
881 cycle. *Journal of Petrology*, submitted.
882

- 883 Herzberg, C.T., and O'Hara, M.J. (1985) Origin of mantle peridotite and komatiite by
884 partial melting. *Geophysical Research Letters*, 12, 541-544.
- 885 Herzberg, C., and Zhang, J. (1996) Melting experiments on anhydrous peridotite KLB-1:
886 Compositions of magmas in the upper mantle and transition zone. *Journal of Geophysical*
887 *Research*, 101, 8271-8295.
- 888
889 Herzberg, C., and Asimow P.D. (2008) Petrology of some oceanic island basalts:
890 PRIMELT2.XLS software for primary magma calculation, *Geochemistry, Geophysics,*
891 *Geosystems* 8, doi:10.1029GC002057.
- 892
893 Herzberg, C., and O'Hara, M.J. (2002) Plume-associated ultramafic magmas of Phanerozoic age.
894 *Journal of Petrology*, 43, 1857-1883.
- 895
896 Herzberg, C., and E. Gazel (2009) Petrological evidence for secular cooling in mantle plumes.
897 *Nature*, 458, 619-622.
- 898
899 Herzberg, C., and Asimow, P.D. (2015) PRIMELT3 MEGA.XLSM software for Primary
900 Magma Calculation: Peridotite Primary Magma MgO Contents from the Liquidus to the Solidus.
901 *Geochemistry, Geophysics, Geosystems*, 16, 563-578, doi:10.1002/2014G005631.
- 902
903 Herzberg, C., Condie, K., and Korenaga, J. (2010) Thermal history of the Earth and its
904 petrological expression. *Earth and Planetary Science Letters*, 292, 79-88.
- 905
906 Herzberg, C., Asimow, P., Ionov, D., Vidito, C., Jackson, M.G., and Geist, D. (2013) Nickel and
907 helium evidence for melt above the core-mantle boundary. *Nature*, 493, 393-397.
- 908
909 Herzberg, C., Cabral, R.A., Jackson, M.D., Vidito, C., Day, J.M.D., and Hauri, E. (2014)
910 Phantom Archean crust in Mangaia hotspot lavas and the meaning of heterogeneous mantle.
911 *Earth and Planetary Science Letters*, 396, 97-106.
- 912
913 Hole, M.J., and Millett, J.M. (2016) Controls of mantle potential temperature and lithospheric
914 thickness on magmatism in the North Atlantic Igneous Province. *Journal of Petrology*, in press.
- 915
916 Ionov, D.A. (2007) Compositional variations and heterogeneity in fertile lithospheric
917 mantle: peridotite xenoliths in basalts from Tariat, Mongolia. *Contributions to Mineralogy and*
918 *Petrology*, 154, 455-477.
- 919
920 Ionov, D.A. (2010) Petrology of mantle wedge lithosphere: New data on supra-subduction zone
921 peridotite xenoliths from the andesitic Avacha volcano, Kamchatka. *Journal of Petrology*, 51,
922 327-361.
- 923
924 Ionov, D.A., and Hofmann, A.W. (2007) Depth of formation of subcontinental off-craton
925 peridotites. *Earth and Planetary Science Letters*, 261, 620-634.

- 926 Ionov, D.A., Ashchepkov, I., and Jagoutz, E. (2005) The provenance of fertile off-craton
927 lithospheric mantle: Sr-Nd isotope and chemical composition of garnet and spinel peridotite
928 xenoliths from Vitim, Siberia. *Chemical Geology*, 217, 41-74.
- 929 Ito, G., and Mahoney, J.J. (2005) Flow and melting of a heterogeneous mantle: 1. Method and
930 importance to the geochemistry of ocean island and mid-ocean ridge basalts. *Earth and Planetary
931 Science Letters*, 230, 29–46.
- 932 Jackson, M.G., Carlson, R.W., Kurz, M.D., Kempton, P.D., Francis, D., and Blusztajn, J. (2010).
933 Evidence for the survival of the oldest terrestrial mantle reservoir. *Nature*, 466, 853-856.
934
- 935 Jenner, F.E., and O'Neill, H. St.C. (2012) Analysis of 60 elements in 616 ocean floor basaltic
936 glasses. *Geochemistry, Geophysics, Geosystems*, 13, Q02005, doi:10.1029/2011GC004009.
937
- 938 Jennings, E.S., and Holland, T.J.B. (2015) A simple thermodynamic model for melting of
939 peridotite in the system NCFMASOCr. *Journal of Petrology*, 56, 869-892.
940
- 941 Jones, J.H. (1984) Temperature and pressure- independent correlations of olivine-liquid partition
942 coefficients and their application to trace element partitioning. *Contributions to Mineralogy and
943 Petrology*, 88, 126-132.
944
- 945 Labrosse, S., Hernlund, J.W., and Coltice, N. (2007) A crystallizing dense magma ocean at the
946 base of the Earth's mantle. *Nature*, 450, 866-869.
947
- 948 Larsen, L.M., and Pedersen, A.K. (2000) Processes in high-Mg, high-T magmas: evidence from
949 olivine, chromite and glass in Palaeogene picrites from West Greenland. *Journal of Petrology*,
950 41, 1071-1098.
951
- 952 Larsen, L.M., and Pedersen, A.K. (2009) Petrology of the Paleocene picrites and flood basalts on
953 Disko and Nuussuaq, West Greenland. *Journal of Petrology*, 50, 1667-1711.
954
- 955 Lee, C.-T., Leeman, W.P., Canil, D., and Li, Z.-X.,A. (2005). Similar V/Sc systematic in MORB
956 and arc basalts: implications for oxygen fugacities of their mantle sources. *Journal of Petrology*,
957 46, 2313-2336.
958
- 959 Leshner, C.E., and Walker, D. (1988) Cumulate maturation and melt migration in a temperature
960 gradient. *Journal of Geophysical Research*, 93, 10295-10311.
961
- 962 Li, C., and Ripley, E.M. (2010) The relative effects of composition and temperature on olivine-
963 liquid Ni partitioning: Statistical deconvolution and implications for petrologic modeling.
964 *Chemical Geology*, 275, 99-104.
965
- 966 Li, J., and Agee, C.B. (2001) The effect of pressure, temperature, oxygen fugacity and
967 composition on partitioning of nickel and cobalt between liquid Fe-Ni-S alloy and liquid silicate:
968 implications for the earth's core formation. *Geochimica et Cosmochimica Acta*, 65, 1821-1832.

- 969
970 Li, X., Kind, R., Priestley, K., Sobolev, S.V., Tilmann, F., Yuan, X. & Weber, M. (2000)
971 Mapping the Hawaiian plume conduit with converted seismic waves. *Nature*, 405, 938-941.
- 972 Li, X. Kind, R. Yuan, X., Wölbern, I. & Hanka, W. (2004) Rejuvenation of the lithosphere by
973 the Hawaiian plume. *Nature*, 427, 827-829.
- 974 Lightfoot, P. C., Hawkesworth, C. J., Olshevsky, K., Green, A., Doherty, W., and Keays, R. R.
975 (1997) Geochemistry of Tertiary tholeiites and picrites from Qeqertarsuaq (Disko Island) and
976 Nuussuaq, West Greenland with implications for the mineral potential of comagmatic intrusions.
977 *Contributions to Mineralogy and Petrology*, 128, 139-163.
- 978
979 Longhi, J., Durand, S.R., and Walker, D. (2010) The pattern of Ni and Co abundances in lunar
980 olivines. *Geochimica et Cosmochimica Acta*, 74, 784-798.
- 981
982 Lupu, R.E., Zahnle, K., Marley, M.S., Schaefer, L., Fegley, B., Morley, C., Cahoy, K., Freedom,
983 R., and Fortney, J.J. (2014) The atmospheres of Earthlike planets after giant impact events. *The*
984 *Astrophysics Journal*, 784, 27. (doi:10.1088/0004-637X/784/1/27).
- 985
986 Mallmann, G., and O'Neill, H.St.C. (2009) The crystal/melt partitioning of V during mantle
987 melting as a function of oxygen fugacity compared with some other elements (Al, P, Ca, Sc, Ti,
988 Cr, Fe, Ga, Y, Zr, and Nb). *Journal of Petrology*, 50, 1765-1794.
- 989
990 Matzen, A.K., Baker, M.B., Beckett, J.R., and Stolper, E.M. (2011) Fe-Mg partitioning between
991 olivine and high-magnesium melts and the nature of Hawaiian parental magmas. *Journal of*
992 *Petrology*, 52, 1243-1263.
- 993
994 Matzen, A.K., Baker, M.B., Beckett, J.R., and Stolper, E.M. (2013) The temperature and
995 pressure dependence of nickel partitioning between olivine and silicate melt. *Journal of*
996 *Petrology*, 54, 2521-2545.
- 997
998 Matzen, A.K., Baker, M.B., Beckett, J.R., Wood, B.J., and Stolper, E.M. (2016) The effect of
999 liquid composition on the partitioning of Ni between olivine and silicate melt. *Contributions to*
1000 *Mineralogy and Petrology*, submitted.
- 1001
1002 McNamara, A.K., Garnero, E.J., and Rost, S. (2010) Tracking deep mantle reservoirs with ultra-
1003 low velocity zones. *Earth and Planetary Science Letters*, 299, 1-9.
- 1004
1005 McKenzie, D. (1984) The generation and compaction of partial melts. *Journal of Petrology*, 25,
1006 713-765.
- 1007
1008 McDonough, W.F., and Sun, S.-s. (1995) The composition of the Earth. *Chemical Geology*, 120,
223-253.

- 1009 Melosh, H. J. (1990). Giant impacts and the thermal state of the early Earth. In Origin of the
1010 Earth (Edited by J. H. Jones and H. E. Newsom) pp. 69-84, Oxford University Press, New York.
1011
- 1012 Morgan, J.P. (2001) Thermodynamics of pressure release melting of a veined plum pudding
1013 mantle. *Geochemistry, Geophysics, Geosystems*, 2, 2000GC000049.
1014
- 1015 Mulyukova, E., Steinberger, B., Dabrowski, M., and Sobolev, S.V. (2015) Survival of LLSVPs
1016 for billions of years in a vigorously convecting mantle: replenishment and destruction of
1017 chemical anomaly. *Journal of Geophysical Research, Solid Earth*, 120,
1018 doi:10.1002/2014JB011688.
1019
- 1020 Mysen, B.O. (2006) Redox equilibria and melt structure: implications for olivine/melt element
1021 partitioning. *Geochimica et Cosmochimica Acta*, 70, 3121-3138.
1022
- 1023 Mysen, B.O. (2007) Partitioning of calcium, magnesium, and transition metals between olivine
1024 and melt governed by the structure of the silicate melt at ambient pressure. *American
1025 Mineralogist*, 92, 844-862.
1026
- 1027 Niu, Y., Wilson, M., Humphreys, E.R., and O'Hara, M.J. (2011) The origin of intra-plate ocean
1028 island basalts (OI B): the lid effect and its geodynamic implications. *Journal of Petrology*, 52,
1029 1443-1468.
1030
- 1031 O'Hara, M.J. (1968) The bearing of phase equilibria studies in synthetic and natural systems on
1032 the origin of basic and ultrabasic rocks. *Earth Science Reviews*, 4, 69-133.
- 1033 O'Hara, M. J. (1977) Geochemical evolution during fractional crystallization of a periodically
1034 refilled magma chamber. *Nature*, 266, 503–507.
1035
- 1036 O'Hara M. J., and Herzberg C. (2002) Interpretation of trace element and isotope features of
1037 basalts: relevance of field relations, petrology, major element data, phase equilibria, and magma
1038 chamber modeling in basalt petrogenesis. *Geochimica et Cosmochimica Acta*, 66, 2167–2191.
1039
- 1040 O'Neill, H.St.C and Jenner, F.E. (2012) The global patten of trace-element distributions in ocean
1041 floor basalts. *Nature* 491, 698-704.
1042
- 1043 Rizo, H., Boyet, M., Blichert-Toft, J., and Rosing, M. (2011) Combined Nd and Hf isotope
1044 evidence for deep-seated source of Isua lavas. *Earth and Planetary Science Letters*, 312, 267–
1045 279.
1046
- 1047 Roeder P.L., and Emslie, R.F. (1970) Olivine-liquid equilibrium. *Contributions to Mineralogy
1048 and Petrology*, 29, 275-289.
1049
- 1050 Pertermann, M., and Hirschmann, M.M. (2003) Anhydrous partial melting experiments on
1051 MORB-like eclogite: phase relations, phase compositions and mineral-melt partitioning of major
1052 elements at 2–3 GPa. *Journal of Petrology*, 44, 2173–2201.

- 1053
1054 Puchtel, I.S., Blichert-Toft, J., Touboul, M., Walker, R.J., Byerly, G.R., Nisbet, E.G., and
1055 Anhauser, C.R. (2013). Insights in early Earth from Barberton komatiites: evidence from
1056 lithophile isotope and trace element systematic. *Geochimica et Cosmochimica Acta*, 108, 63-90.
1057
1058 Putirka, K. (2016) Cooling rates for Earth, Moon, Mars and Vesta, and new models for oxygen
1059 fugacity, ferric-ferrous ratios, olivine-liquid Fe-Mg exchange, and mantle potential temperatures.
1060 *American Mineralogist*, in press.
1061
1062 Putirka, K., Ryerson, F.J., Perfit, M., and Ridley, W.I. (2011) Mineralogy and composition of the
1063 oceanic mantle. *Journal of Petrology*, 52, 279-313.
1064
1065 Rubie, D., Frost, D.J., Mann, U., Asahara, Y., Nimmo, F., Tsuno, K., Kegler, P., and Palme, H.
1066 (2011) Heterogeneous accretion, composition, and core-mantle differentiation of the Earth. *Earth
1067 and Planetary Science Letters*, 301, 31-42 (2011).
- 1068 Saunders, A.D., Fitton, J.G., Kerr, A.C., Norry, M.J., and Kent, R.W. (1997) The North Atlantic
1069 Igneous Province (edited by J.J. Mahoney and M.J. Coffin) *Geophysical Monograph Series*, 100,
1070 pp. 45-93, AGU, Washington DC.
1071
1072 Shorttle, O., MacLennan, J. (2011) Compositional trends of Icelandic basalts: implications for
1073 short-length scale lithological heterogeneity in mantle plumes. *Geochemistry, Geophysics,
1074 Geosystems*, 12, Q11008. <http://dx.doi.org/10.1029/2011GC003748>.
1075
1076 Shorttle, O., MacLennan, J., and Lambart, S. (2014) Quantifying lithological variability in the
1077 mantle. *Earth and Planetary Science Letters*, 195, 24-40.
1078
1079 Siebert, J., Badro, J., Antonangeli, D., and Ryerson, F.J. (2012) Metal-silicate partitioning of Ni
1080 and Co in a deep magma ocean. *Earth and Planetary Science Letters*, 321, 189-197.
1081
1082 Sobolev, A.V., Hofmann, A.W., Sobolev, S.V., and Nikogosian I.K. (2005) An olivine-free
1083 mantle source of Hawaiian shield basalts. *Nature*, 434, 590-597.
1084
1085 Sobolev, A.V., Hofmann, A.W., Kuzmin, D.V., Yaxley, G.M., Arndt, N.T., Chung, S.-L.,
1086 Danyushevsky, L.V., Elliott, T., Frey, F.A., Garcia, M.O., Gurenko, A.A., Kamenetsky, V.S.,
1087 Kerr, A.C., Krivolutskaya, N.A., Matvienkov, V.V., Nikogosian, I.K., Rocholl, A., Sigurdsson,
1088 I.A., Sushchevskaya, N.M., and Teklay M. (2007) The amount of recycled crust in sources of
1089 mantle-derived melts. *Science*, 316, 412-417.
1090
1091 Sobolev, A.V., Asafov, E.V., Gurenko, A.A., Arndt, N.T., Batanova, V.G., Portnyagin, M.V.,
1092 Garbe-Schönberg, D., and Krasheninnikov, S.P. (2016) Komatiites reveal a hydrous Archaean
1093 deep-mantle reservoir. *Nature*, 531, 628-632.
1094
1095 Solomatov, V. (2015) Magma oceans and primordial mantle differentiation. In *Treatise on
1096 Geophysics*, (Shubert, G. Ed.), 2nd edition, 9, 81-104.

- 1097
1098 Starkey, N.A., Stuart, F.M., Ellam, R.M., Fitton, J.G., Basu, S., and Larsen, L.M. (2009) Helium
1099 isotopes in early Iceland plume picrites: Constraints on the composition of high $^3\text{He}/^4\text{He}$ mantle.
1100 Earth and Planetary Science Letters, 277, 91-100.
1101
1102 Straub, S. M., LaGatta, A. B., Martin-Del Pozzo, A. L. and Langmuir, C. H. (2008) Evidence
1103 from high Ni olivines for a hybridized peridotite/pyroxenite source for orogenic andesites from
1104 the central Mexican Volcanic Belt. *Geochemistry, Geophysics, Geosystems*, 9, Q03007,
1105 doi:10.1029/2007GC001583.
1106
1107 Stuart, F.M., Lass-Evans, S., Fitton, J.G., and Ellam, R.M. (2003) High $^3\text{He}/^4\text{He}$ ratios in picritic
1108 basalts from Baffin Island and the role of a mixed reservoir in mantle plumes. *Nature*, 424, 57-
1109 59.
1110
1111 Takahashi, E. (1978) Partitioning of Ni^{2+} , Co^{2+} , Fe^{2+} , Mn^{2+} , and Mg^{2+} between olivine and
1112 silicate melts: compositional dependence of partition coefficient. *Geochimica et Cosmochimica*
1113 *Acta*, 42, 1829-1844.
1114
1115 Tonks, W.B., and H. J. Melosh, H.J. (1990) The physics of crystal settling and suspension in a
1116 turbulent magma ocean. In *Origin of the Earth* (edited by J. H. Jones and H. E. Newsom), 151-
1117 174, Oxford University Press, New York.
1118
1119 Tonks, W.B., and Melosh, H.J. (1993) Magma ocean formation due to giant impacts. *Journal of*
1120 *Geophysical Research*, 98, 5319–5333.
1121
1122 Trela, J., Vidito, C., Gazel, E., Herzberg, C., Class, C., Whalen, W., Jicha, B., Bizimis, M., and
1123 Alvarado, G.E. (2015) Recycled crust in the Galápagos plume source at 70 Ma: implications for
1124 plume evolution. *Earth and Planetary Science Letters* 425, 268-277.
1125
1126 Walker, D. (2005) Core-mantle chemical issues. *Canadian Mineralogist*, 43, 553-1564.
1127
1128 Walter, M.J. (1998) Melting of garnet peridotite and the origin of komatiite and depleted
1129 lithosphere. *Journal of Petrology*, 39, 29-60.
1130
1131 Walter, M.J., and Trønnes, R.G. (2004) Early Earth differentiation. *Earth and Planetary Science*
1132 *Letters*, 225, 253-269.
1133
1134 Walter, M.J., and Cottrell, E. (2013) Assessing uncertainty in geochemical models for core
1135 formation in Earth. *Earth and Planetary Science Letters*, 365, 165-176.
1136
1137 Wang, Z., and Gaetani, G.A. (2008) Partitioning of Ni between olivine and siliceous eclogite
1138 partial melt: experimental constraints on the mantle source of Hawaiian basalts. *Contributions to*
1139 *Mineralogy and Petrology*, 156, 661-678.
1140

1141 Weis, D., Garcia, M.O., Rhodes, J.M., Jellinek, M., and Scoates, J.S. (2011) Role of deep mantle
1142 in generating the compositional asymmetry of the Hawaiian mantle plume. *Nature Geoscience*,
1143 4, 831-838.
1144
1145 Yaxley, G.M., and Green, D.H. (1998). Reactions between eclogite and peridotite: mantle
1146 refertilisation by subduction of oceanic crust. *Schweiz. Mineral. Petrogr. Mitt.*, 78, 243–255.
1147
1148 Zhang, J., and Herzberg, C.T. (1994) Melting experiments on anhydrous peridotite KLB-1 from
1149 5.0 to 22.5 GPa. *Journal of Geophysical Research*, 99, 17,729-17,742.
1150
1151
1152
1153
1154
1155
1156
1157
1158
1159
1160
1161
1162
1163
1164
1165
1166
1167
1168
1169
1170
1171
1172
1173
1174
1175
1176
1177
1178
1179
1180
1181
1182
1183
1184
1185

1186 **FIGURE 1.** Experimentally measured $D_{\text{NiO}}^{\text{Ol/L}}$ and its model prediction (Beattie et al., 1991;
1187 Matzen et al., 2013), where $D_{\text{NiO}}^{\text{Ol/L}}$ is expressed on a weight% basis (i.e., wt% NiO in
1188 olivine/wt% NiO in melt). For Beattie et al. (1991), weight % D = mole% \times 1.135. For Matzen
1189 et al. (2013), weight % D = mole% \times 0.5318. The database consists of 284 individual
1190 experimental olivine-melt pairs, and is given in Table A2 of the Appendix.

1191
1192 **FIGURE 2.** Model nickel contents of peridotite-source primary magmas and their equilibrium
1193 olivines. Primary magma compositions are solutions to the equations for batch and accumulated
1194 fractional melting (i.e., AFM) using an initial source composition having 0.25% NiO (i.e., 1964
1195 ppm Ni). Equilibrium olivine compositions at pressures of melt generation are computed from
1196 Beattie-Jones, Toplis, and Herzberg-O'Hara parameterizations. Please see text for details. a)
1197 Blue lines are tie lines connecting olivine composition to liquids in equilibrium with olivine only
1198 (i.e., L+Ol). Red line labeled solidus are Ni and MgO contents of melts on the solidus; black
1199 space between red and green lines are Ni contents of primary melts for the assemblages
1200 (L+Ol \pm Opx \pm Cpx \pm Sp \pm Gr). b) Effects of peridotite composition and melting mechanism on the
1201 Ni and MgO contents of melts. Green squares are off-craton peridotite xenolith data from Asia
1202 (Ionov 2007; 2010; Ionov and Hofmann 2007; Ionov et al. 2005).

1203
1204 **FIGURE 3.** Nickel contents and Mg numbers (i.e., mole% 100MgO/(MgO + FeO)) for model
1205 and measured olivines. a) black and red forms are olivine compositions in equilibrium with batch
1206 and accumulated fractional melts, respectively, using the Beattie-Jones parameterization. Gray
1207 shaded form with $\pm 1\sigma$ represents the 1 *Absolute* RMSE uncertainty in olivine Ni content that is
1208 propagated from uncertainties in experimental data (Table A2 and Herzberg et al. 2013). Black
1209 cross is olivine composition at the liquidus for near-total melting. Mg number is mole%
1210 100MgO/(MgO + FeO). Blue form bounds olivine compositions calculated by Straub et al.
1211 (2008); Mg numbers for olivines appropriate for near-solidus and near-liquid melts are \sim 89.5
1212 and 96.0. b) comparison of model olivines of batch primary magmas with experimentally-
1213 measured olivines (Herzberg and Zhang 1996; Table A3); MgO and FeO contents of olivines
1214 distributed between the liquidus and solidus were illustrated in Figures 4, 5, and 6 in Herzberg
1215 and Zhang (1996).

1216
1217 **FIGURE 4.** Nickel contents and Mg numbers (i.e., mole% 100MgO/(MgO + FeO)) for model
1218 olivines compared with 215 measured olivines in mantle peridotite (Herzberg et al. 2013). a)
1219 model olivines in equilibrium with accumulated fractional melts in the mantle. b) model olivines
1220 in equilibrium with batch melts in the mantle.

1221
1222 **FIGURE 5.** Nickel contents and Mg numbers (i.e., mole% 100MgO/(MgO + FeO)) of model
1223 olivines compared with measured olivines from Archean komatiites and Makapuu-stage Koolau
1224 volcano, Hawaii (Sobolev et al. 2007). Melting is assumed to be accumulated fractional
1225 (Herzberg 2004), and olivine crystallizes from primary magmas at the surface; this is more
1226 clearly shown and described in Figure 6. Green lines are model CLD for primary magmas
1227 having the MgO contents in the 20 to 30% range; note the model olivine CLD is concaved
1228 toward the Mg number axis, in good agreement with measured olivines.

1229

1230 **FIGURE 6.** Nickel contents and Mg numbers (i.e., mole% 100MgO/(MgO + FeO)) of model
1231 and measured olivines. Model olivine compositions for a Siqueiros MORB primary magma is
1232 given in Herzberg and Asimow (2015; MgO = 11.56%), and is likely representative of all
1233 MORB. Open cross is olivine in equilibrium with the primary magma having a Ni content as
1234 described by equation (10). Olivine crystal line of descent CLD is compared with measured
1235 olivines from the Siqueiros fracture zone (Table A4) and South East Indian Ridge (Sobolev et al.
1236 2007). The CLD calculation was made with PRIMELT3 (Herzberg and Asimow 2015) for
1237 fractionation of olivine, and MELTS (Ghioro and Sack 1995) for fractionation of olivine +
1238 plagioclase and olivine + plagioclase + clinopyroxene at 200 bars, an oxygen fugacity defined by
1239 the QFM buffer and with 0.2% H₂O. Olivine compositions were computed using Fe-Mg
1240 partitioning from Toplis (1995). Ni partition coefficients used: olivine-melt from Beattie et al.
1241 (1991); plagioclase-melt = 0; clinopyroxene-melt = 0.25 D_{NiO}^{Ol/L} (Herzberg et al. 2013). It is
1242 notable that MELTS is not sufficiently calibrated for Ni, and also underestimates olivine Mg
1243 number by ~ 2%; we use the MELTS CLD and compute olivine Mg number from Toplis (1995).
1244 Note also that model olivines at high pressure have slightly lower Mg numbers than those at the
1245 surface, and this is accommodated by higher K_D in the models of Herzberg and O'Hara (2002)
1246 and Toplis (2005). Therefore, olivines that crystallize from a primary magma at the surface
1247 (black form) will have higher Mg numbers than residual olivines at mantle depths (red form),
1248 and their range of compositions is highly restricted.

1249
1250 **FIGURE 7.** Nickel contents and Mg numbers (i.e., mole% 100MgO/(MgO + FeO)) of model
1251 olivines compared with measured olivines from Theistareykir volcano, Iceland and Makapuu-
1252 stage Koolau volcano, Hawaii (Sobolev et al. 2007). All model olivines are assumed to
1253 crystallize at 1 atmosphere. Open cross is the composition of olivine that would crystallize from
1254 the primary magma containing 17.6% MgO and Ni as given by equation (10); this is a
1255 PRIMELT3 primary magma solution (Herzberg and Asimow 2015) for Theistareyker lava
1256 sample TH03xrf (Shorttle et al. 2014). Black lines labeled L+Ol, L+Ol+Plag, and
1257 L+Ol+Plag+Cpx describe the compositions of olivine that would crystallize along the liquid line
1258 of descent (i.e., CLD) at 1 atmosphere, anhydrous and at an oxygen fugacity defined by the QFM
1259 buffer. Arrow labeled P describes the effect of pressure on the L+Ol+Plag+Cpx CLD. The CLD
1260 calculation was made with PRIMELT3 (Herzberg and Asimow 2015) and MELTS (Ghioro and
1261 Sack 1995) as described in Figure 6b for MORB, but for 0% H₂O. Red lines are compositions of
1262 olivines that would crystallize from magmas that form by mixing of the primary magma with
1263 derivative magmas along the LLD. White circles are Ni contents of olivine calculated using the
1264 method of Matzen et al. (2013; their experiments combined with a compilation of literature data;
1265 2975 ppm Ni in olivine in residual peridotite). In all cases olivine is assumed to crystallize near
1266 the surface. White circles are for the following 5 conditions of melt segregation in the mantle,
1267 from lowest to highest Ni contents: 1) 0 GPa & 1399°C, 2) 0.5 GPa 1426°C, 3) 1 GPa &
1268 1451°C, 4) 1.5 GPa & 1476°C, 5) 1504°C & P = 2.1 GPa (Hole and Millet 2016). In all cases,
1269 these temperatures T at pressures P are calculated from equations 12 and 13 in Herzberg &
1270 Asimow (2015). Melting pressures refer to final melting of the mantle region, not pressures at
1271 which melting initiates.

1272
1273 **FIGURE 8.** Nickel contents and Mg numbers (i.e., mole% 100MgO/(MgO + FeO)) of model
1274 olivines compared with measured olivines from West Greenland and Baffin Island (Table A5).

1275 Model olivines are those that crystallize from primary magmas at the surface; melting
1276 mechanism is accumulated fractional melting. Olivines that crystallize from primary magmas
1277 have been calculated using the Beattie-Jones model and peridotite sources having 1964
1278 (McDonough and Sun 1995) and 2360 ppm Ni. Crystal line of descent and mixing lines are from
1279 Figure 7. White circles are calculated Ni contents of olivine using the method of Matzen et al.
1280 (2013; their experiments combined with a compilation of literature data; 2975 ppm Ni in olivine
1281 in residual peridotite; primary magma MgO = 20.0%) for the following conditions: high Ni using
1282 melting $T = 1580^{\circ}\text{C}$ & $P = 2.8$ GPa (Herzberg and Asimow 2015; Hole and Millet 2016) and
1283 crystallization $T = 1444^{\circ}\text{C}$ & 1 atmosphere (Herzberg & Asimow 2015); low Ni using melting T
1284 $= 1549^{\circ}\text{C}$ & $P = 2.1$ GPa and crystallization $T = 1444^{\circ}\text{C}$ & 1 atmosphere (Herzberg & Asimow
1285 2015; Hole and Millet 2016). Melting pressures refer to final melting of the mantle region, not
1286 pressures at which melting initiates.

1287 .
1288 **FIGURE 9.** Fe/Mn and Mg numbers of model olivines compared with measured olivines. Data
1289 from the South East India Rise, Theistareykir, Hawaii are from Sobolev et al. (2007); data from
1290 West Greenland and Baffin Island are from this work (Table A5). Black form is model olivines
1291 of a peridotite source having 0.13% MnO (Herzberg 2011). Measured olivines from Iceland are
1292 consistent with model olivines of a peridotite source having Mg numbers > 90 and Fe/Mn in the
1293 62 to 66 range. Fractional crystallization of olivine does not change Fe/Mn very much, but
1294 clinopyroxene fractionation can elevate Fe/Mn (Herzberg et al. 2014; Trela et al. 2015).
1295 Olivines from Hawaii have Fe/Mn that are higher than those expected of peridotite-source
1296 melting, owing possibly to retention of Mn in a garnet pyroxenite source, and elevation of
1297 primary melt Fe/Mn from which the olivines crystallized (Herzberg 2011).

1298
1299 **FIGURE 10.** Histograms of V/Sc for West Greenland and Baffin Island lavas that have been
1300 filtered with Nd and Sr isotope ratios to exclude samples that have experienced crustal
1301 contamination. West Greenland lavas generally have higher V/Sc, consistent with deeper
1302 melting of garnet peridotite.

1303
1304 **FIGURE 11.** Model cobalt contents of peridotite-source primary magmas and their equilibrium
1305 olivines. Primary magma compositions are solutions to the equations for batch and accumulated
1306 fractional melting (i.e., AFM) using an initial source composition having 105 ppm Co
1307 (McDonough and Sun 1995). Equilibrium olivine compositions at pressures of melt generation
1308 and 1 atmosphere are computed from Beattie-Jones, Toplis, and Herzberg-O'Hara
1309 parameterizations (please see text for details). Gray shaded form in panel b) is the 1σ uncertainty
1310 in olivine composition for AFM partial melts arising from application of the Beattie et al. (1991)
1311 parameterization (i.e., equation 5) to experimental data (Herd et al. 2009; Longhi et al. 2010;
1312 Mysen 2006; 2007; Takahashi 1978).

1313
1314 **FIGURE 12.** Cobalt contents and Mg numbers (i.e., mole% $100\text{MgO}/(\text{MgO} + \text{FeO})$) of model
1315 and measured olivines from mantle peridotite. Model olivines are those in equilibrium with
1316 primary magmas of accumulated fractional melting. Measured olivines are those for San Carlos
1317 olivine (Table A6; Sobolev et al. 2007), and olivines in peridotite from orogenic and xenoliths
1318 occurrences (De Hoog et al. 2010).

1319

1320 **FIGURE 13.** Model cobalt contents of peridotite-source primary magmas, their equilibrium
1321 olivines compared with observed cobalt contents of MORB glasses and olivines. Model results
1322 are from Figure 11. a) Measured MORB glasses are from Jenner and O'Neill (2012). b)
1323 Measured olivines are from the Siqueiros transform of the East Pacific Rise (Table A4 and
1324 Sobolev et al. 2007) and Knipovich ridge (Sobolev et al., 2007).

1325
1326 **FIGURE 14.** Model cobalt contents of peridotite-source primary magmas, their equilibrium
1327 olivines compared with observed cobalt contents of whole rocks and olivines from West
1328 Greenland and Baffin Island. Model results are from Figure 11. a) Measured whole rocks are
1329 from Larsen and Pedersen (2009; red symbols). b) Measured olivines are from Baffin Island
1330 (i.e., BI sample numbers) and West Greenland (i.e., WG sample numbers) (Table A5).

1331
1332

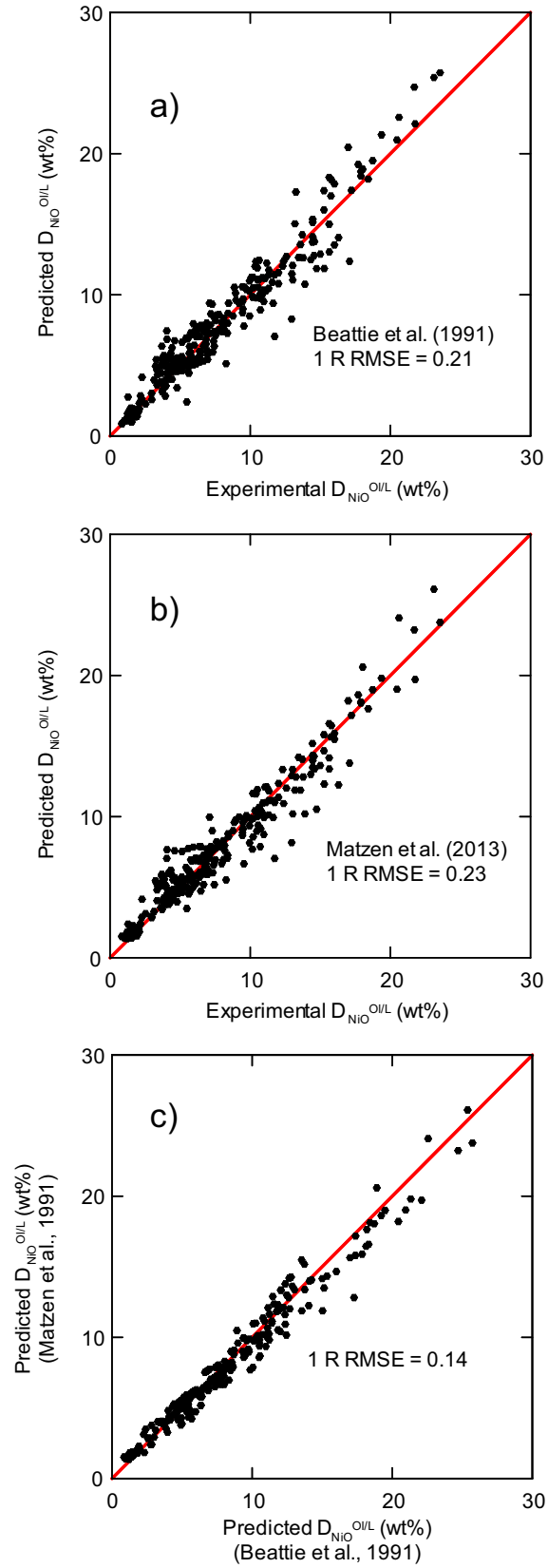


Figure 1

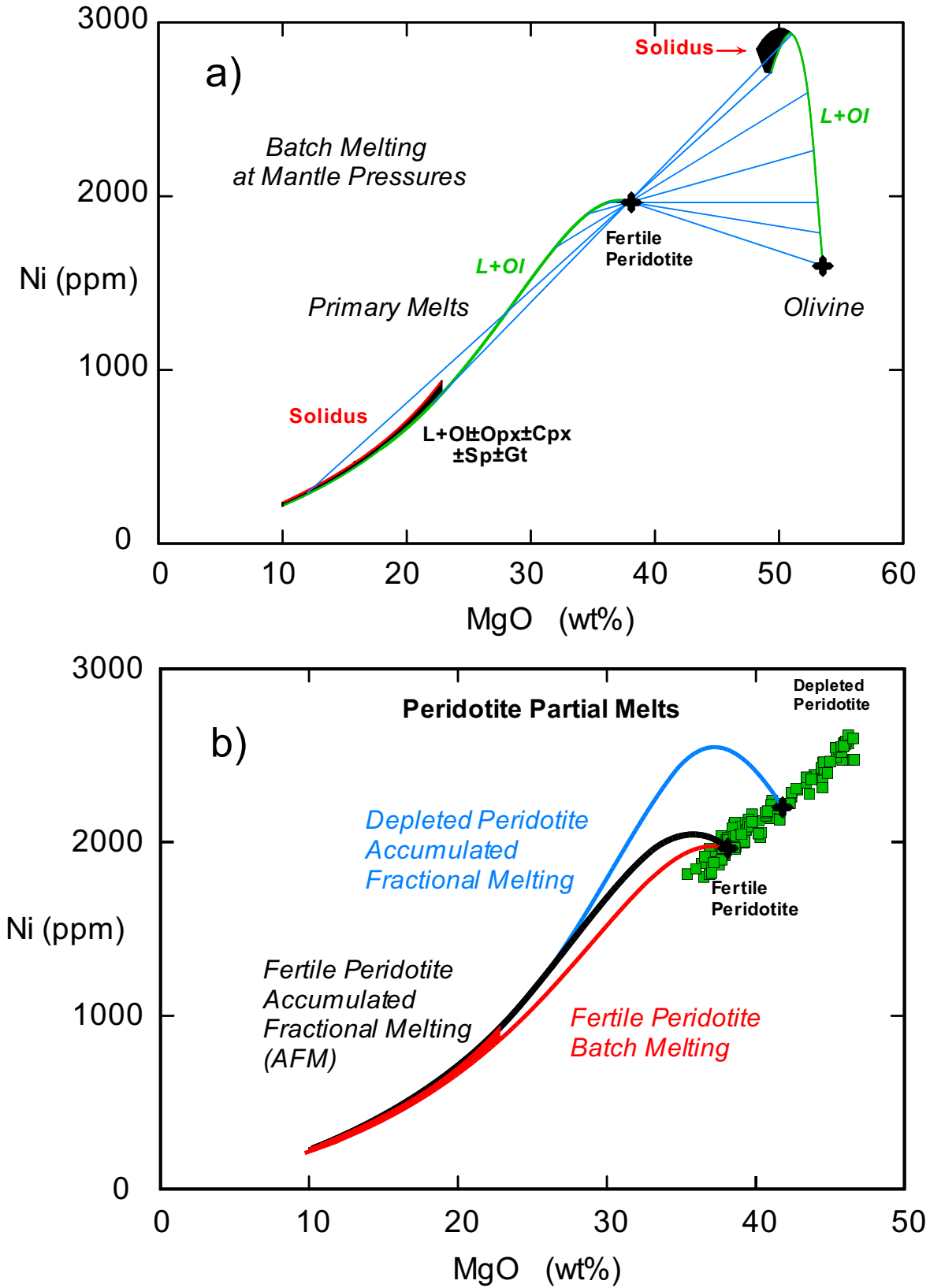


Figure 2

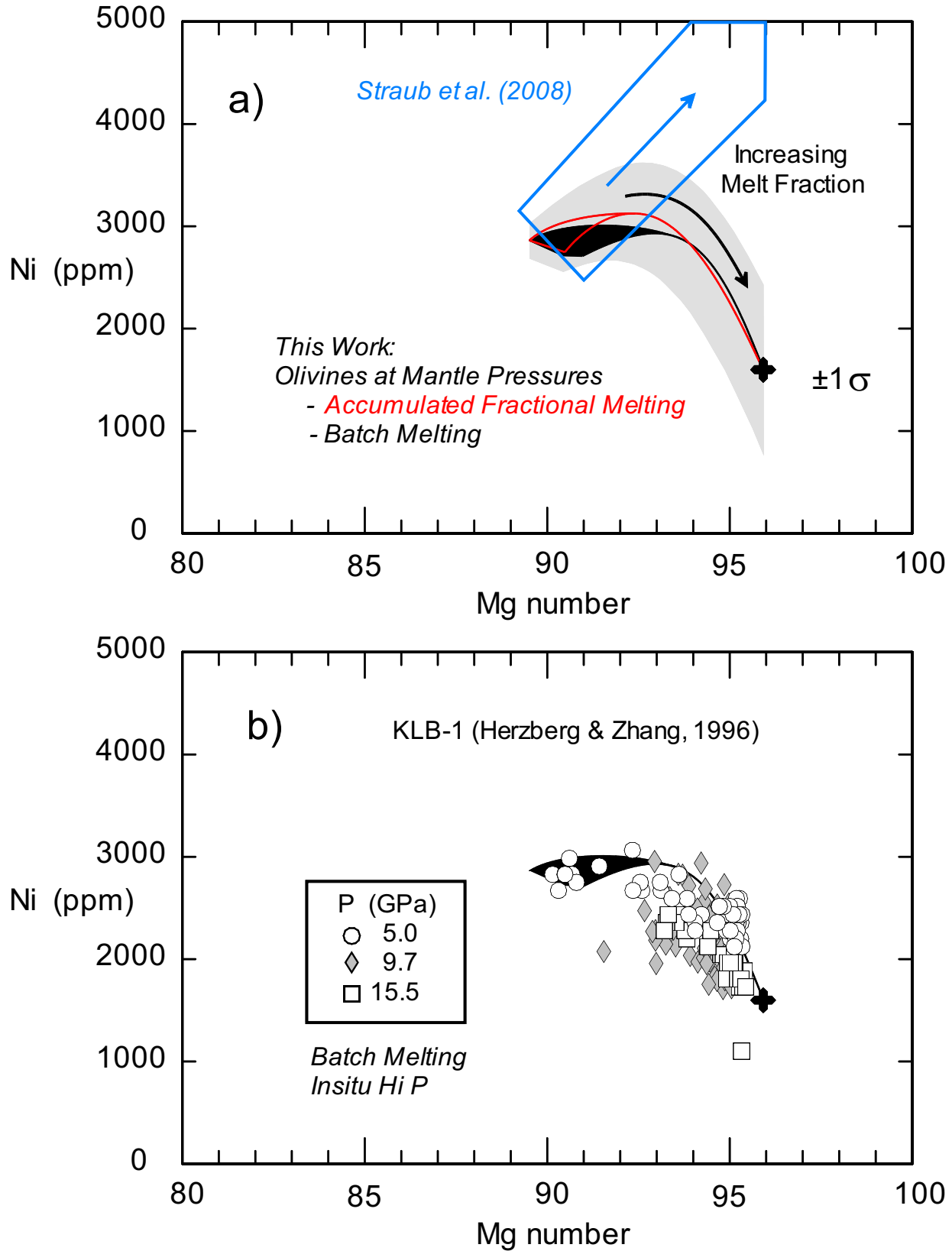


Figure 3

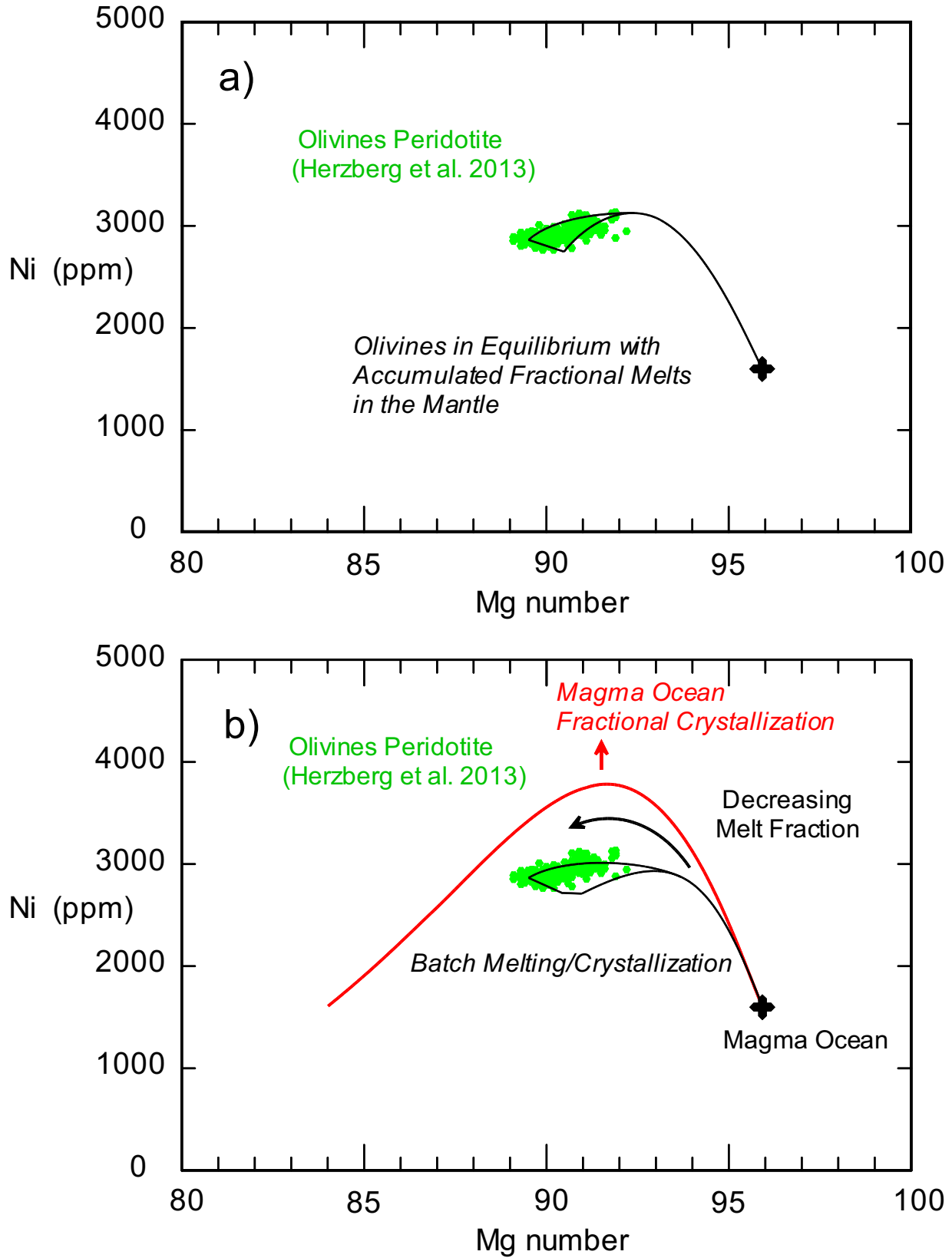


Figure 4

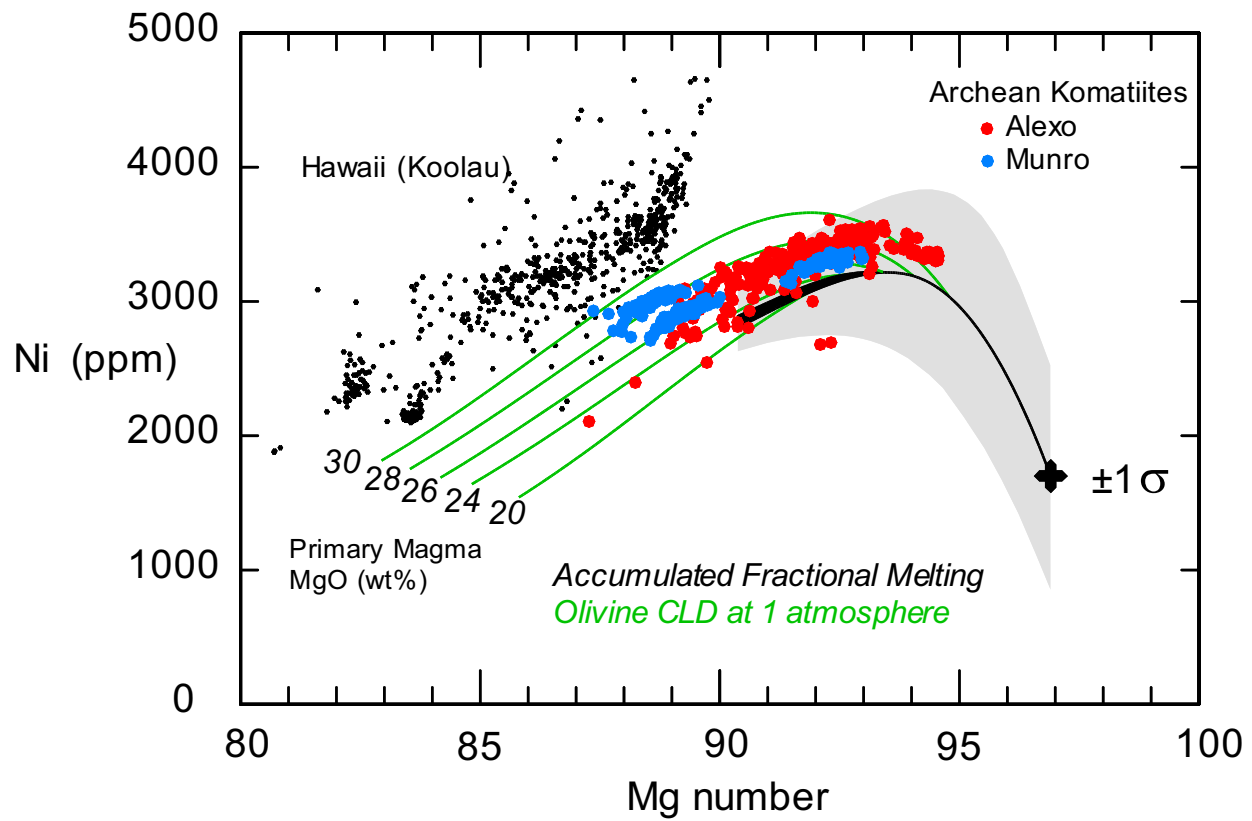


Figure 5

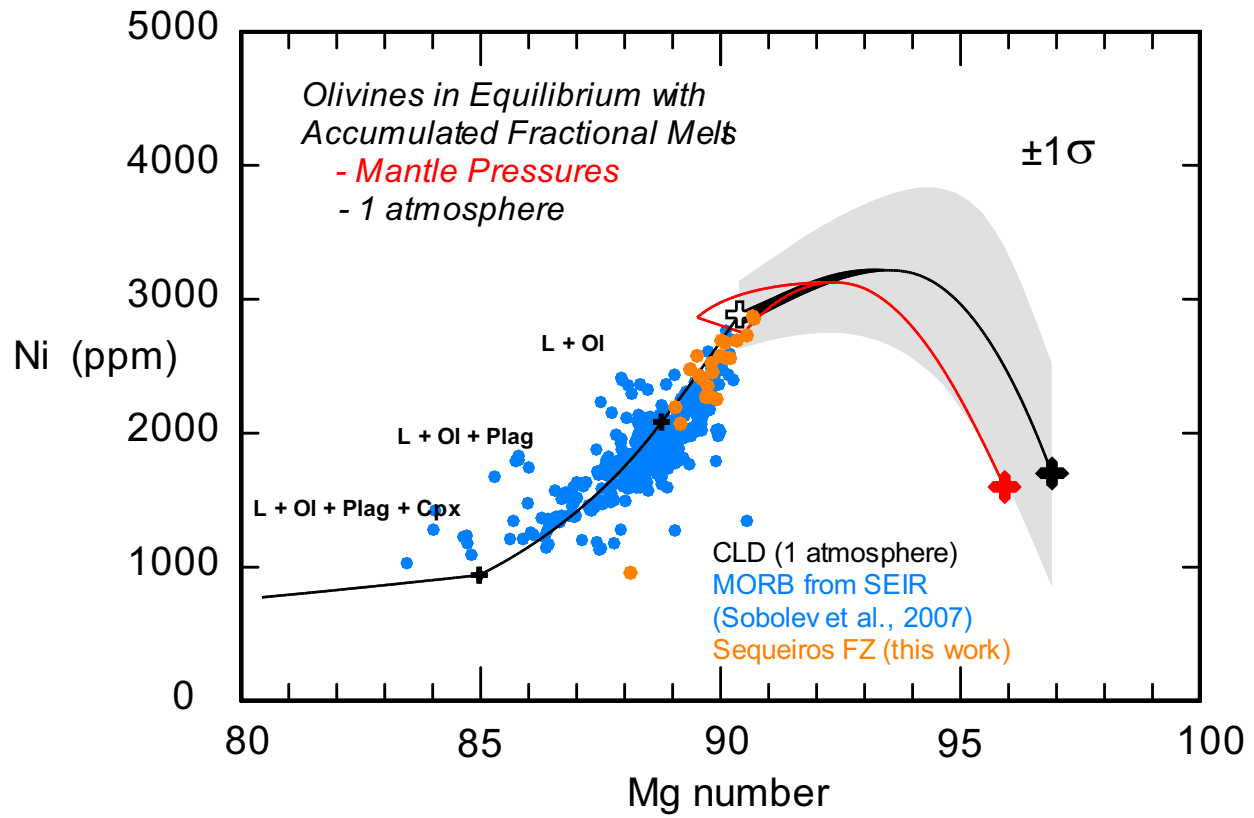


Figure 6

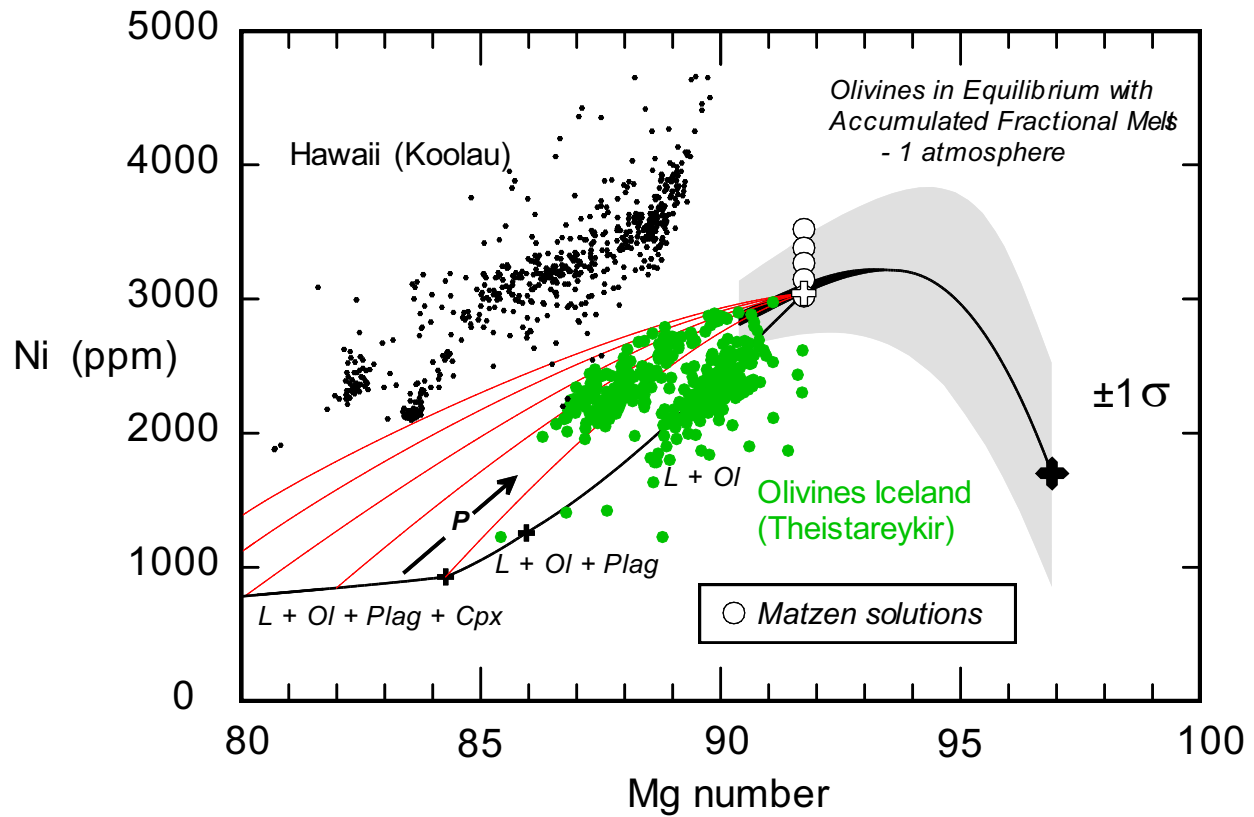


Figure 7

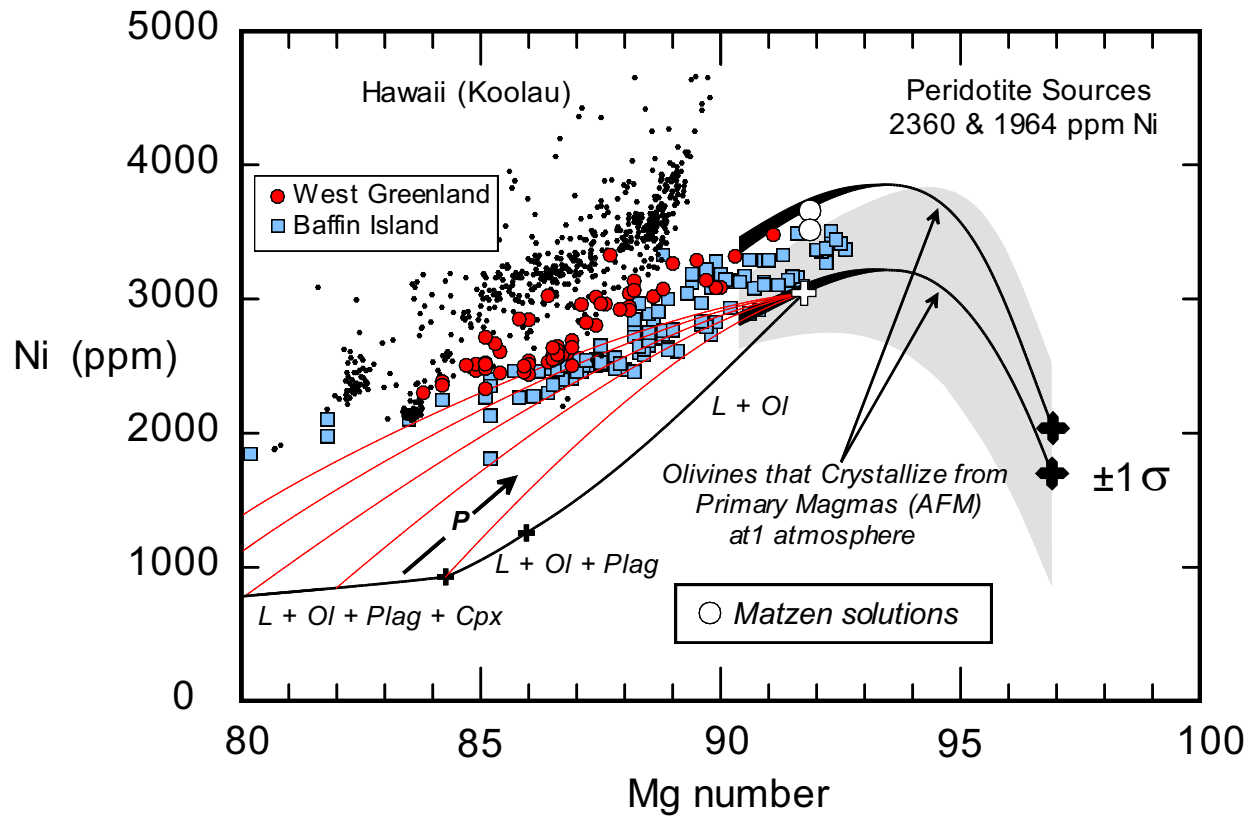


Figure 8

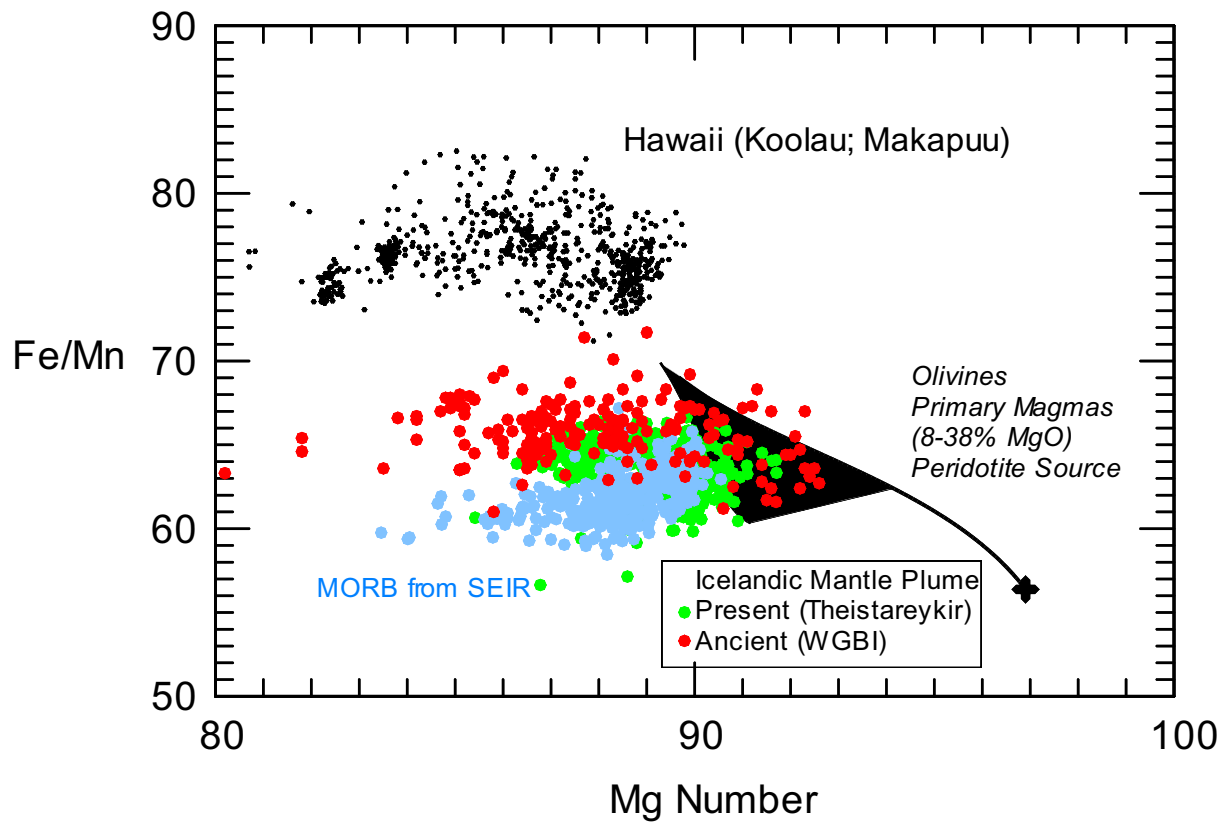


Figure 9

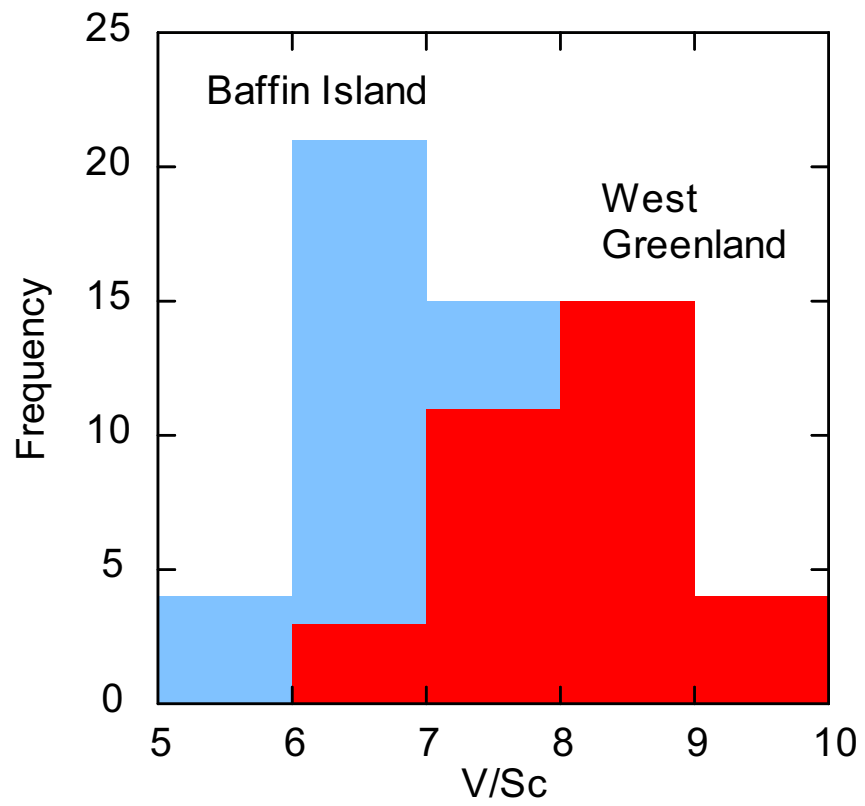


Figure 10

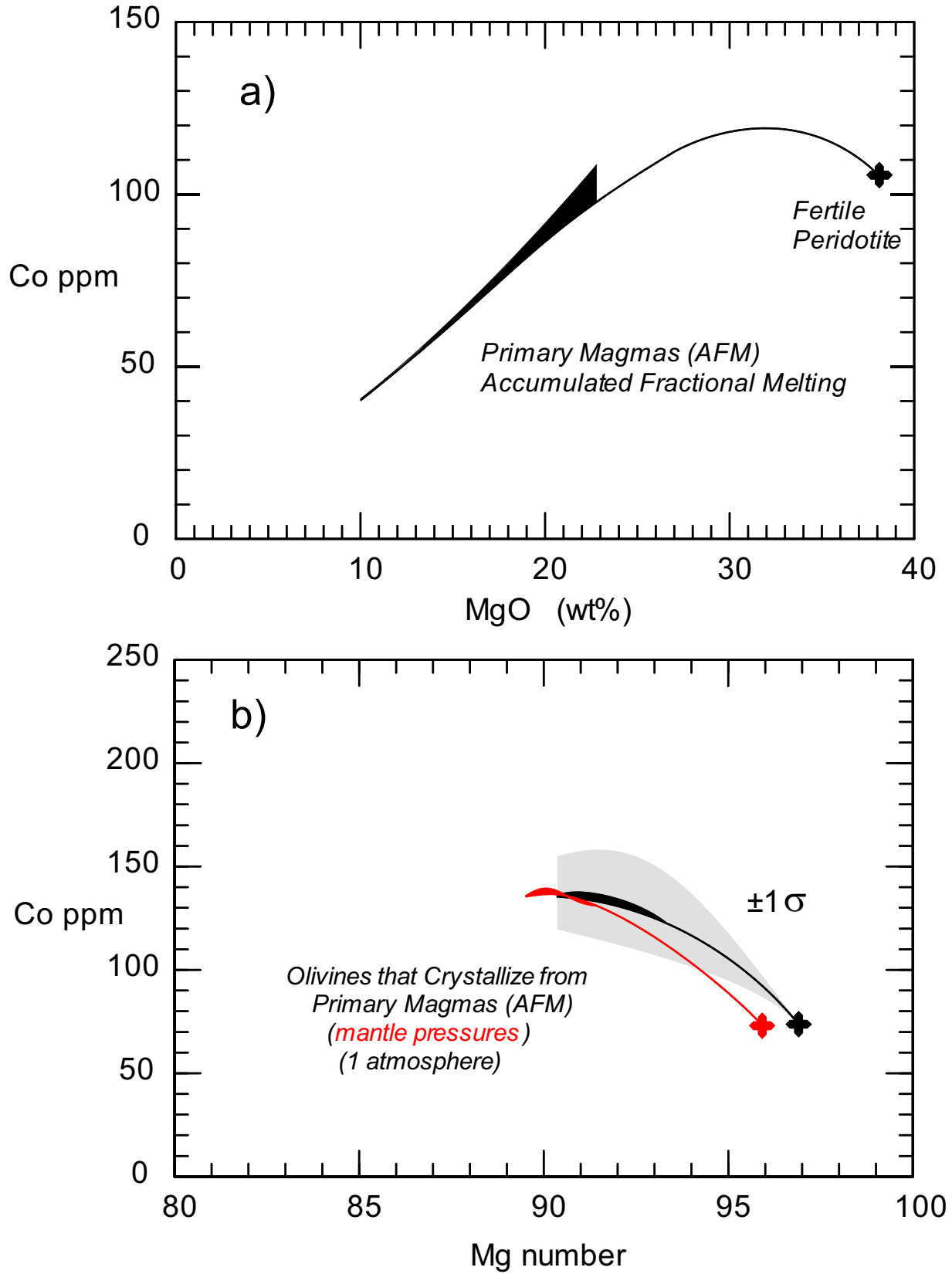


Figure 11

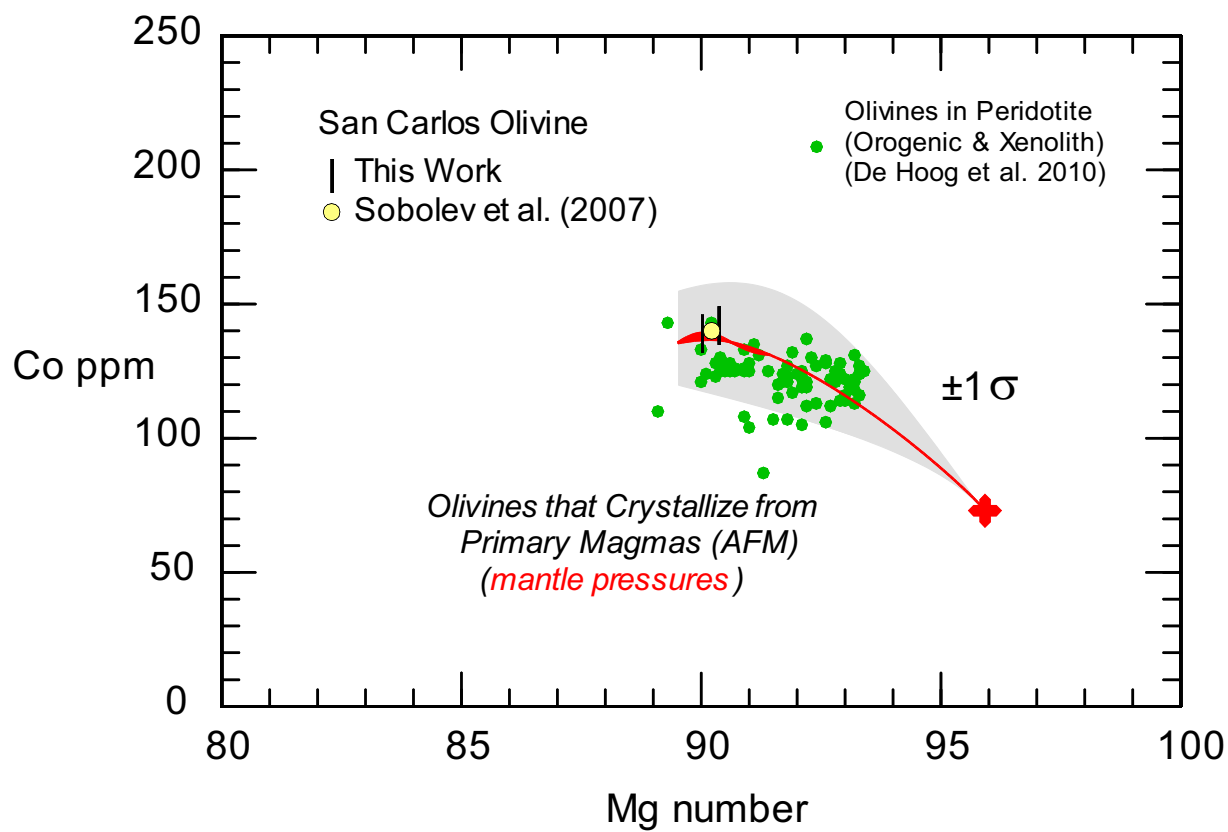


Figure 12

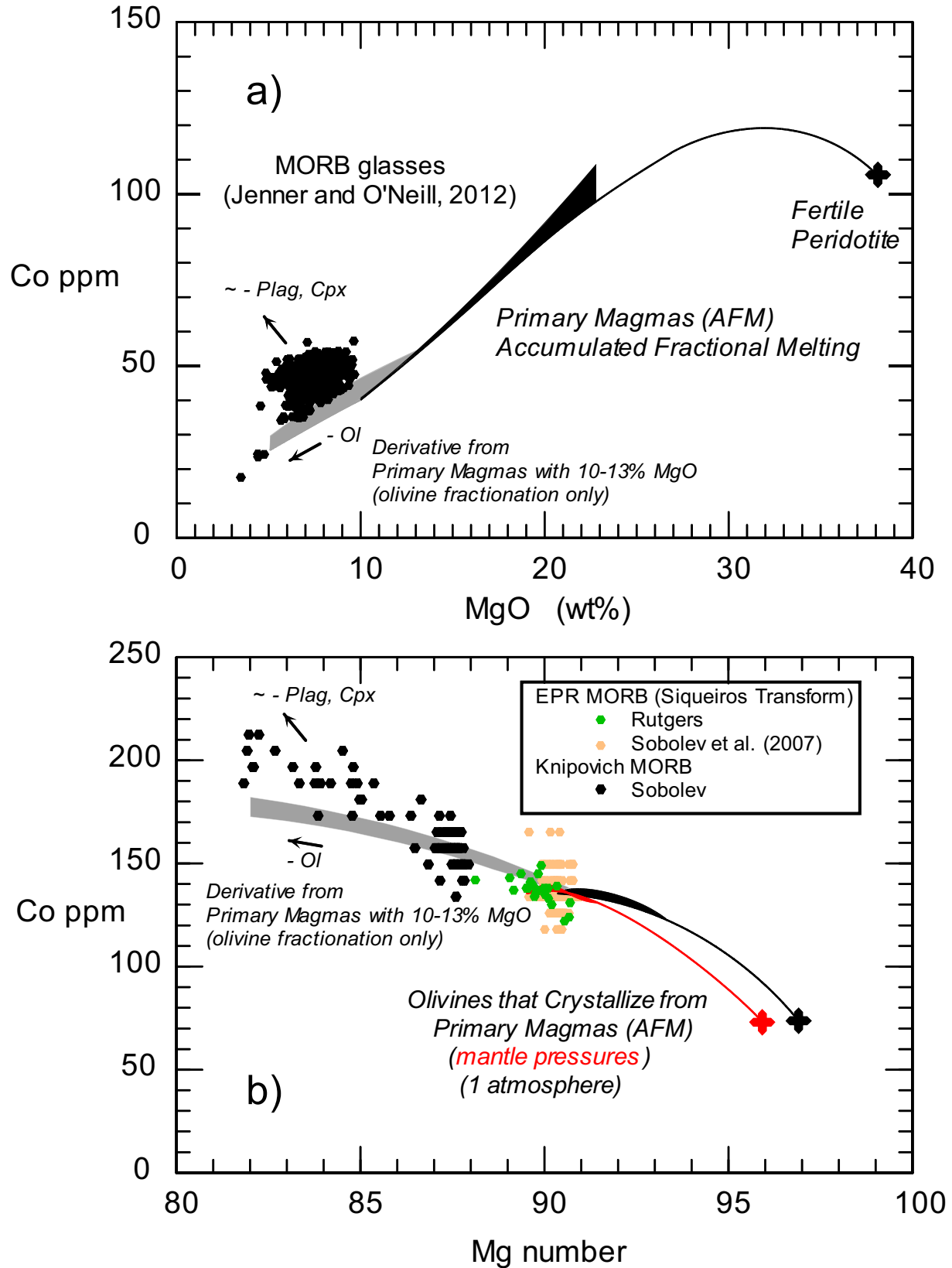


Figure 13

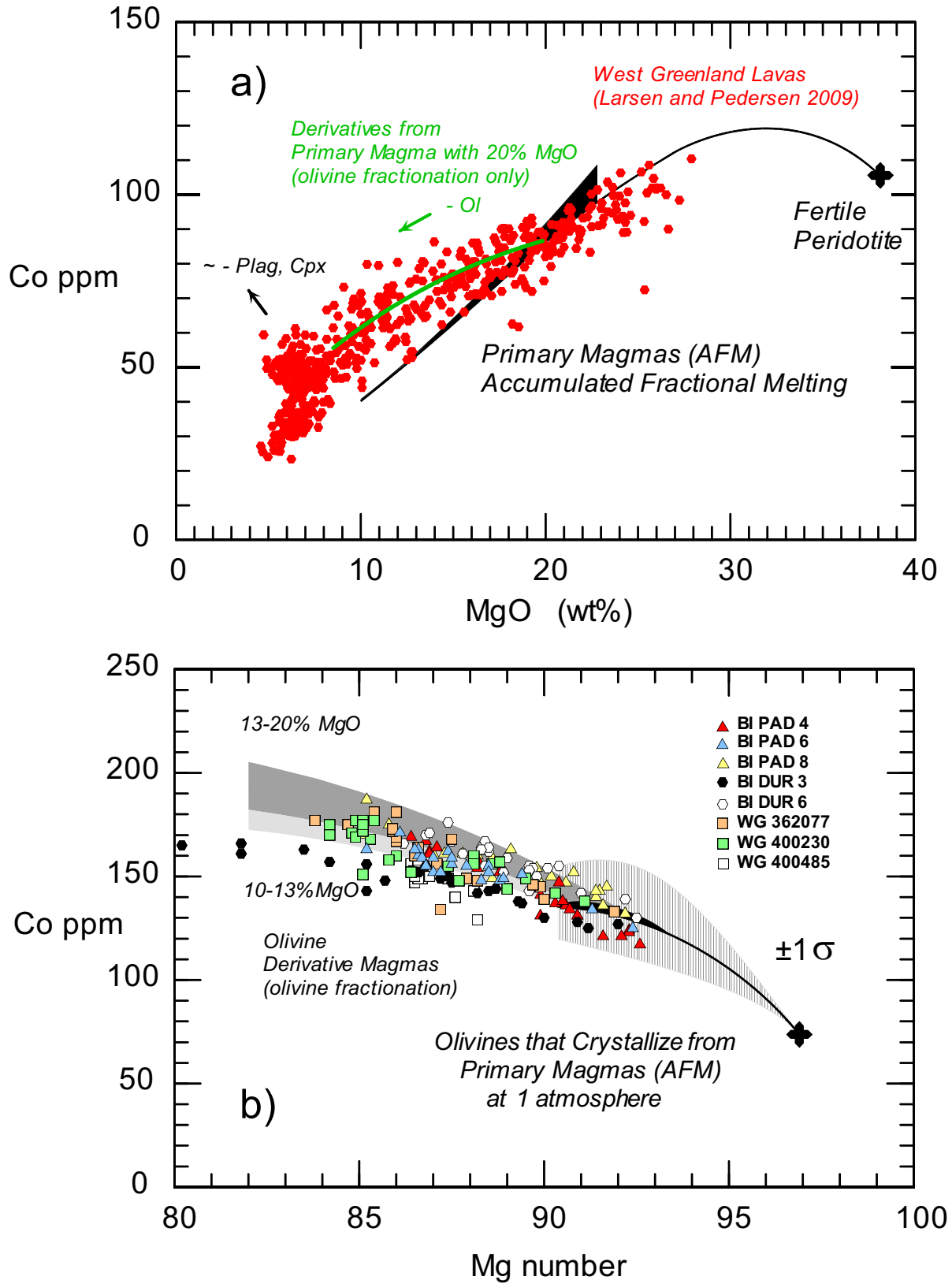


Figure 14